



Leibniz-Institut für
Astrophysik Potsdam

ISC proudly presents:

4. AIP-Jamboree, January 23, 2015

The rules of the game:

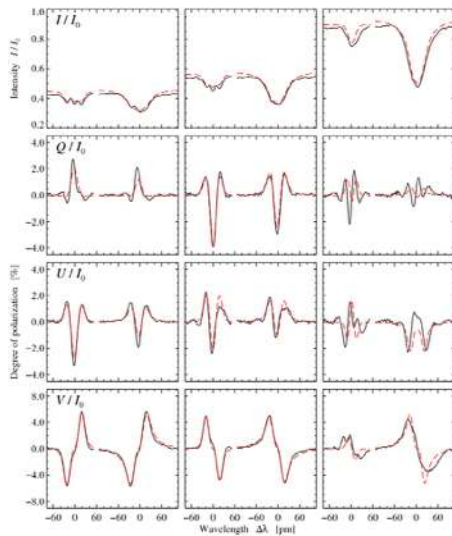
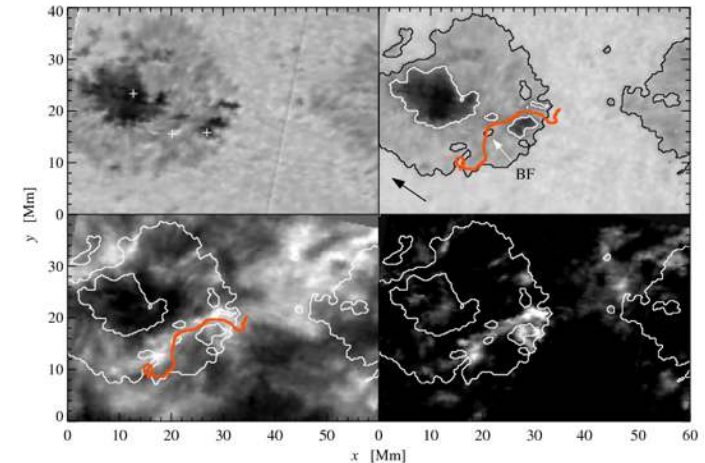
- 2 minutes (2 slides)*
- Present yourself and your work*
- Get to know the colleagues*

Horst Balthasar

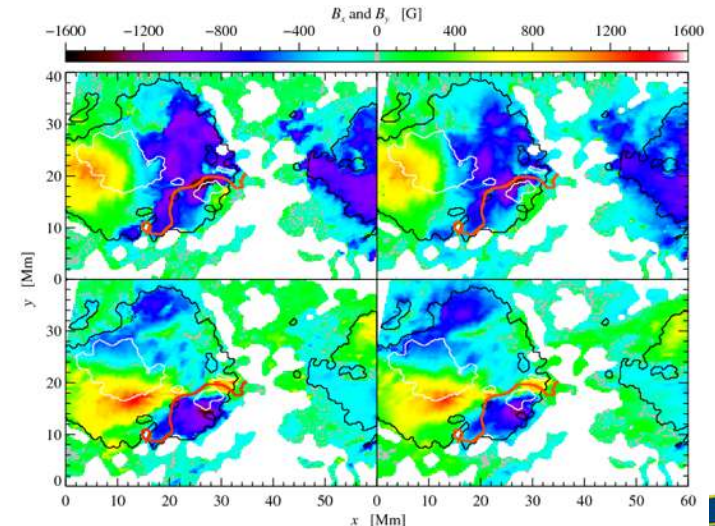
Solar Optical

Magnetic field in sunspots

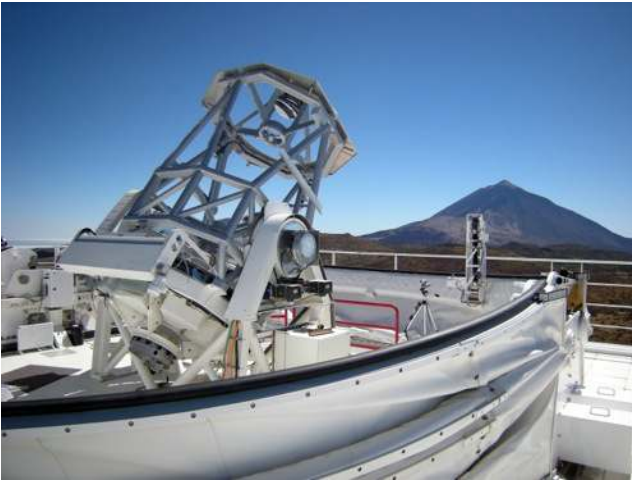
- Measure the Stokes profiles
- run an inversion code
(e.g. Stokes Inversion based on response functions, SIR)
- Solve magnetic ambiguity to get the magnetic vector field
- Example: delta spot NOAA 11504



A&A 562, L6 (2014)
ASP CS 489, 39 (2014)



Instrumental polarization of the GREGOR-telescope

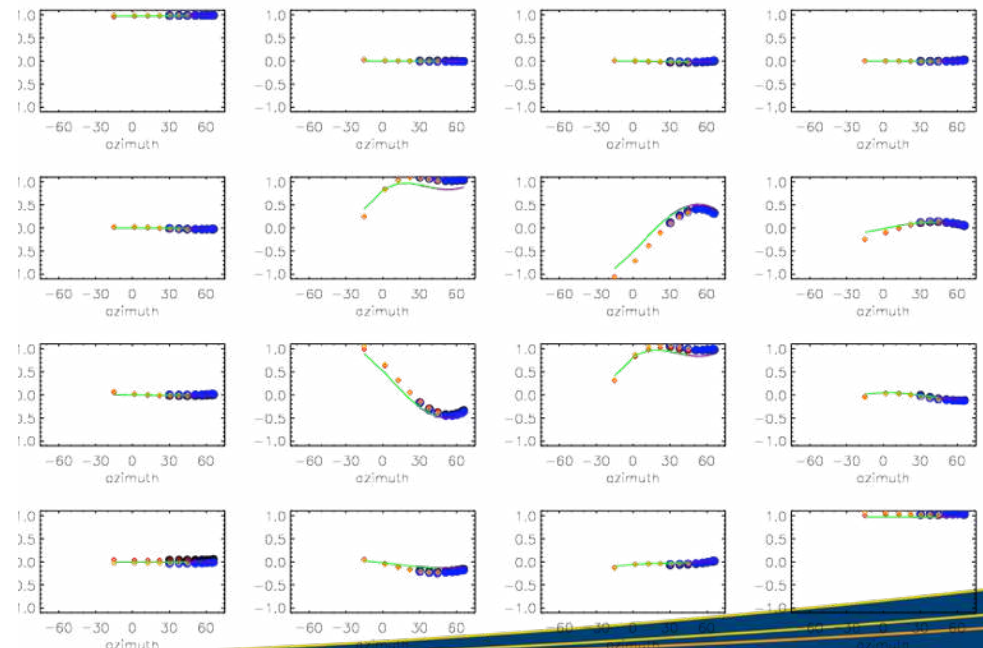
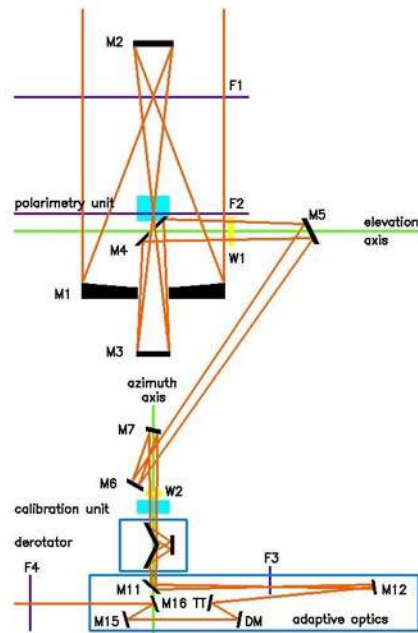


Alt-azimuthal mount — mirrors rotate against each other

Polarimetric modulation inside the post-focus instruments

Need for a telescope model to calibrate data

Hofmann et al., AN 333, 854 (2012)

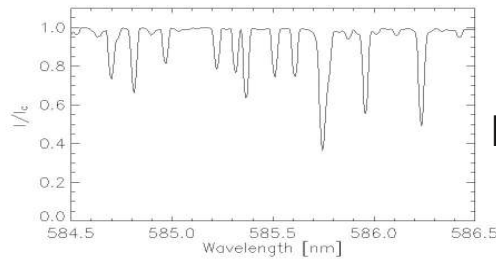
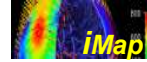


Thorsten Carroll

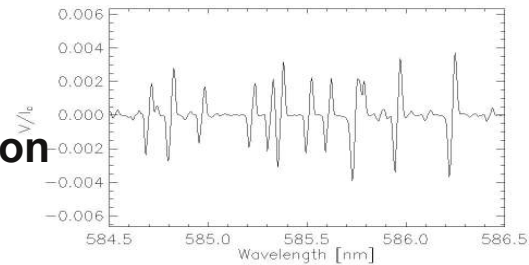
Stellar physics

Mapping Stellar Surfaces - Zeeman-Doppler Imaging

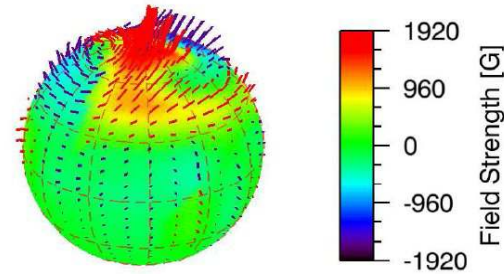
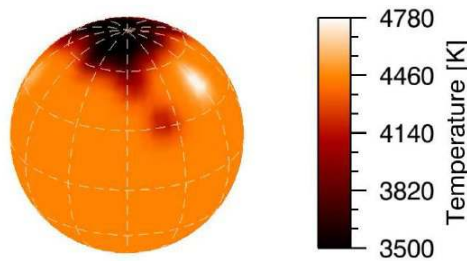
Thorsten A. Carroll – Stellar Physics and Stellar Activity



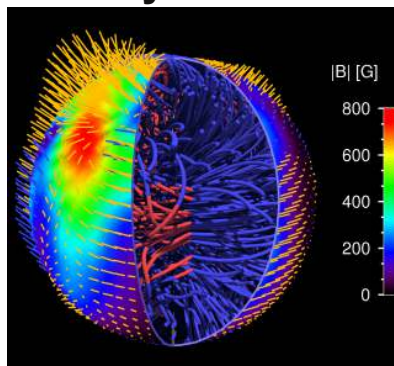
**Spectrum
Intensity & Polarization**



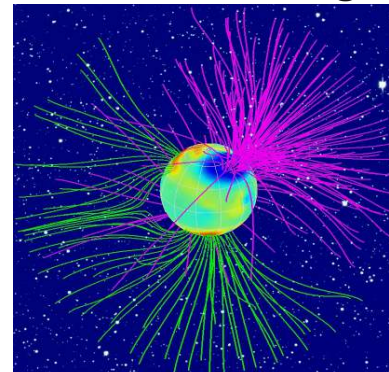
Reconstructed temperature & magnetic field distribution



Comparison with theory Stellar interior



Magnetic field extrapolation Stellar environment



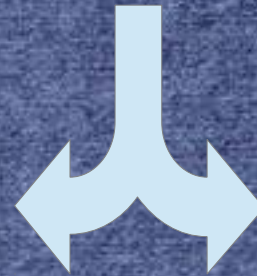
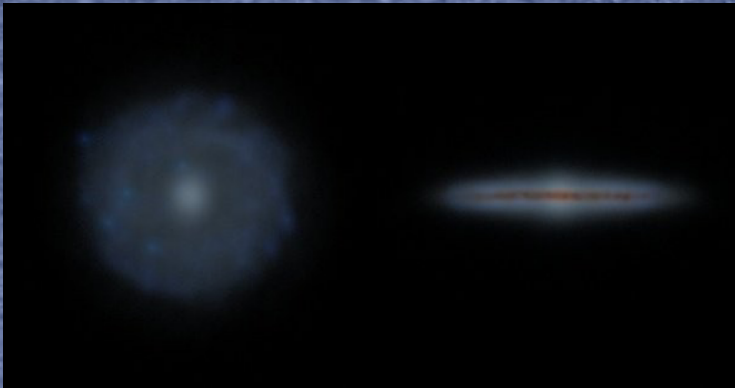
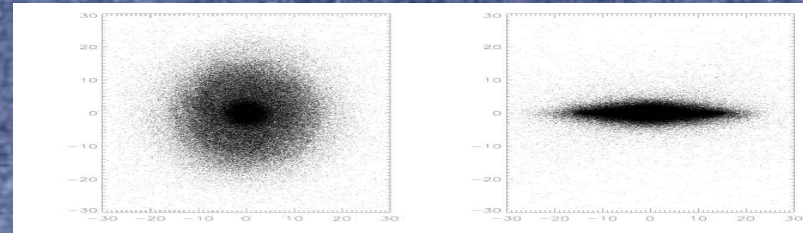
Giovanni Guidi

Cosmology

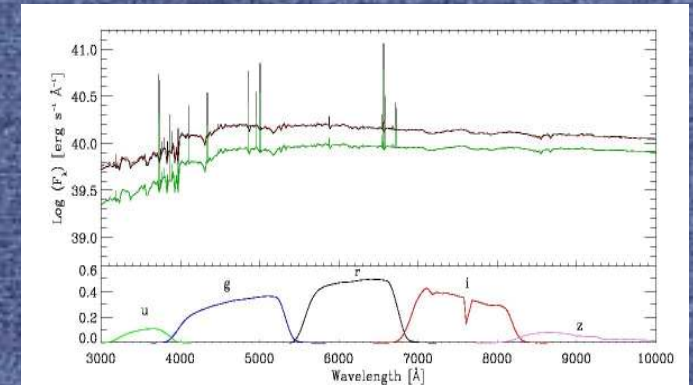
Galaxy formation in cosmological context:

Dr. C. Scannapieco (supervisor), Dr. P. Creasey,
P. Poulhazan (Ph.D.), me (Ph.D.)

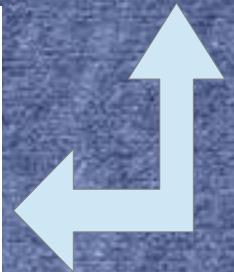
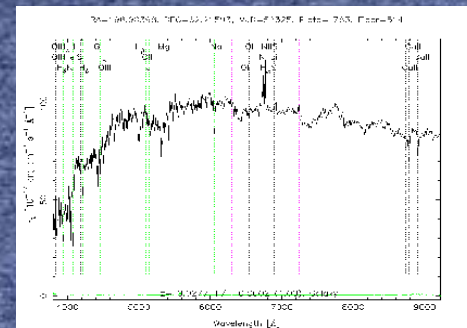
SIMULATION



MOCK DATA

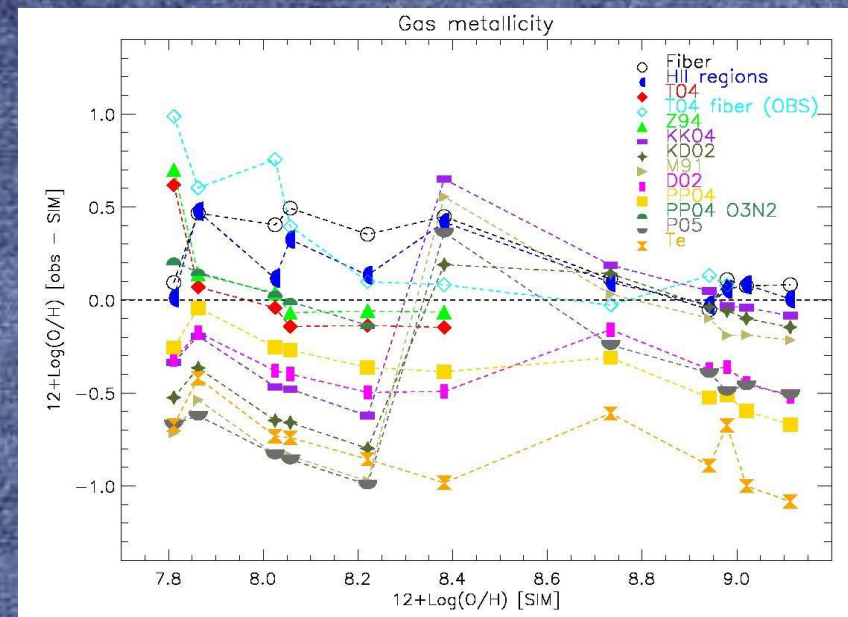
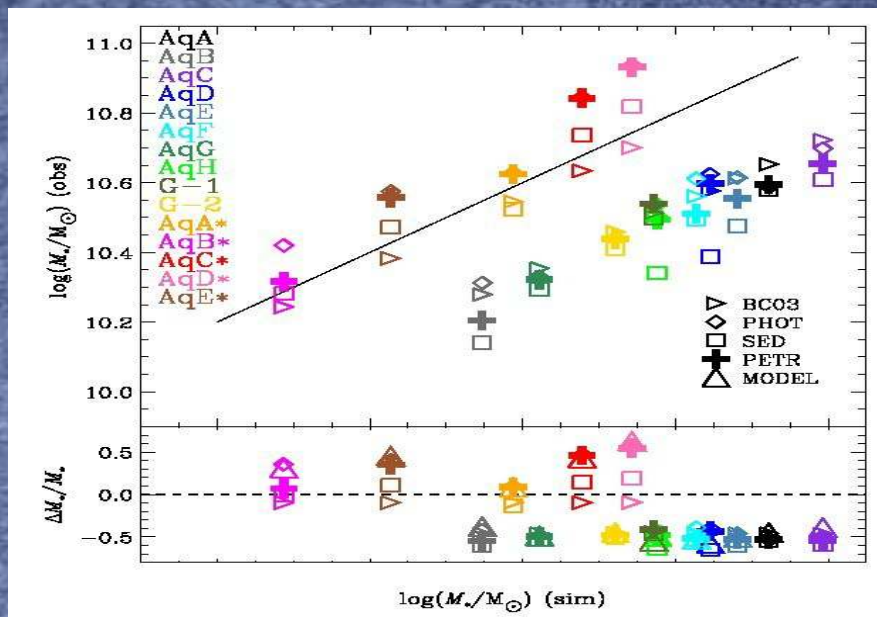
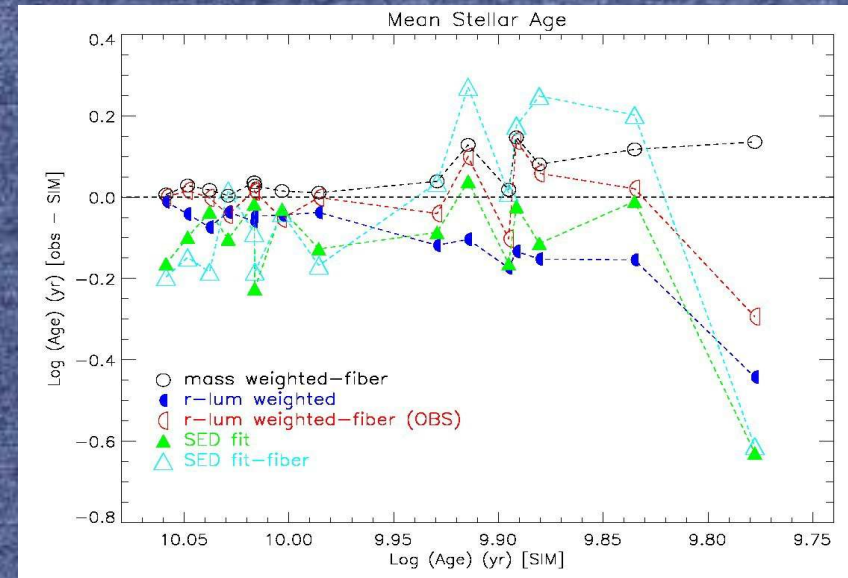


REAL DATA



Derive galaxy properties in “observational” way

- Compare simulations and observations in an unbiased manner
- Test the model of galaxy formation
- Test observer's algorithms



Eloy Luis Hernandez Anguizola

innoFSPEC

Background

- Dipl.-Ing. Electrical Engineering
- Experimental setups and data analysis



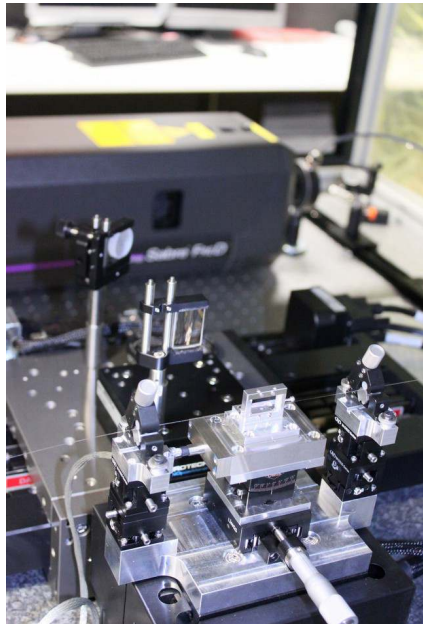
At AIP

- April 2014 part of InnoFSPEC
- October 2014 PhD → supervision: Prof. Martin M. Roth
 - Physical Properties of Optical Fibers used for Spectroscopy in the Optical and NIR at Extremely Large Telescopes



Projects

- PRAXIS: OH suppressed fibre spectrograph
 - IR Detector H2RG



- Fibre Bragg Grating (FBGs)

Arto Järvinen

Robotics



- UoO MSc 2002

- GRB010222

- NOT 2003-2004

- Student support astronomer
 - ALFOSC, FIES, StanCam

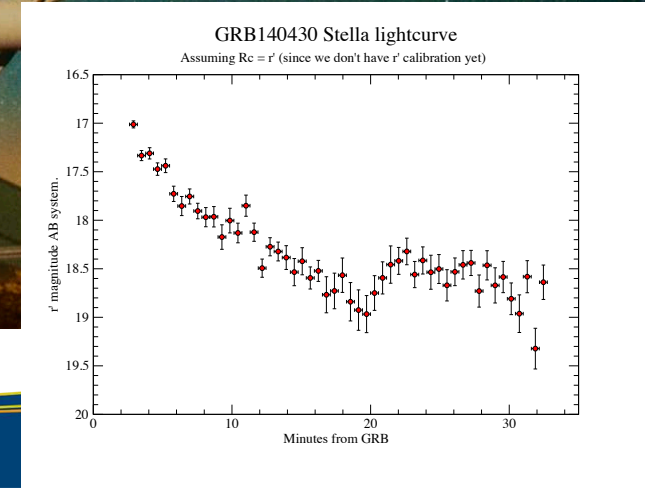
- ESO 2004-2005

- FORS photometry precision

- AIP 2006...

- STELLA, RoboTel, PEPsi, Gregor@night





Roelof de Jong

Milky Way and Local Volume

Galaxy formation and evolution

- Roelof de Jong
 - PhD Groningen -> Durham -> Steward Obs. -> STScI -> AIP (July 2009)
 - Head Milky Way and the Local Volume section
 - We study any galaxy that can be resolved into stars
 - PI GHOSTS & 4MOST

Galaxy

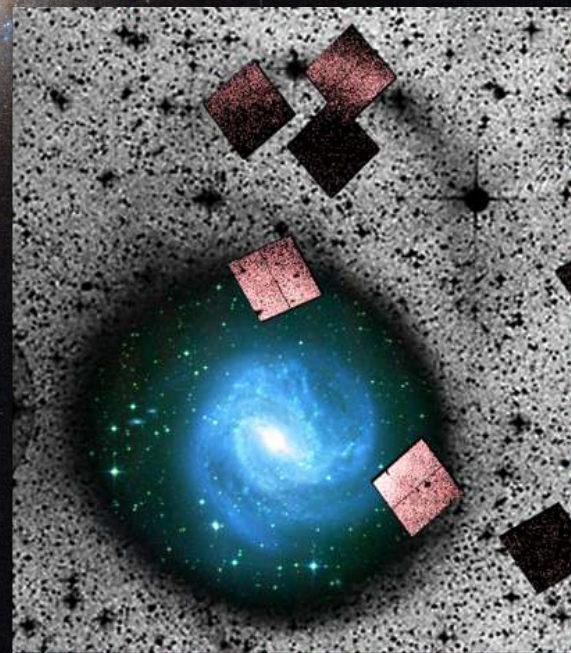
Halos

Outer Disks

Substructure

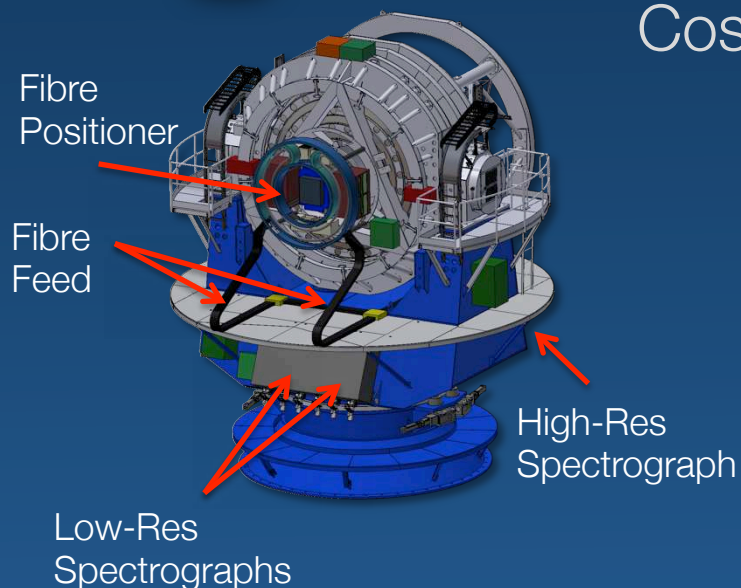
Thick Disks

Star Clusters





4MOST – 4m Multi-Object Spectroscopic Telescope



Cosmology & Large Scale Structure (Euclid)
Galaxy Evolution & Dark Matter
High Energy Universe (eROSITA)
Milky Way Formation & Evolution (Gaia)
Exo-planet host stars (PLATO)

Open to anyone at the AIP!

VISTA telescope

www.4MOST.eu



Thiago Junqueira

Milky Way and Local Volume

Jamboree - AIP

A little bit of my past

Thiago C. Junqueira

Where I'm from



But not from Rio



I'm from São Paulo



Where I studied

I did my bachelor in physics at
State University of São Paulo
2003 – 2006



Then my master and PhD at
University of São Paulo
2007 – 2013



Now here I am

Postdoc since nov – 2013



Support by



How I felt
during the
word cup



Jamboree - AIP

I'm working ...

...in the Milky Way group
more directly with

Cristina
Chiappini



Ivan
Minchev



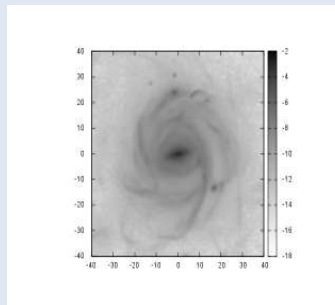
About my work

My work topics are:



Structure of Milky Way:

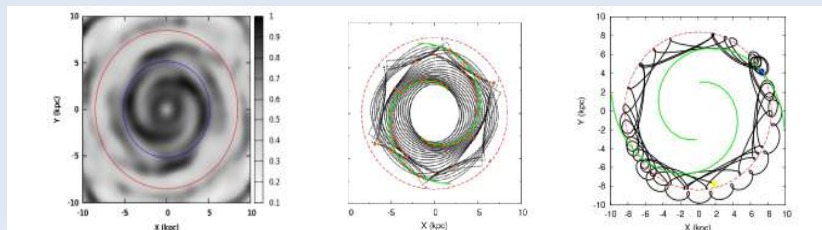
Focus more
on spiral arms
and stellar dynamics



Chemodynamical
evolution

Simulated galaxy
from
Martig et al. (2012)

Stellar dynamics



Interest on the
following surveys:

RAVE

SEGUE

APOGEE

GES

and in the
near future

Gaia

4MOST

WEAVE.

Francisco Kitaura
Schwarzschild Fellow

The bottom of the slide features several thin, wavy, light-colored lines that sweep across the width of the image, creating a subtle decorative border.

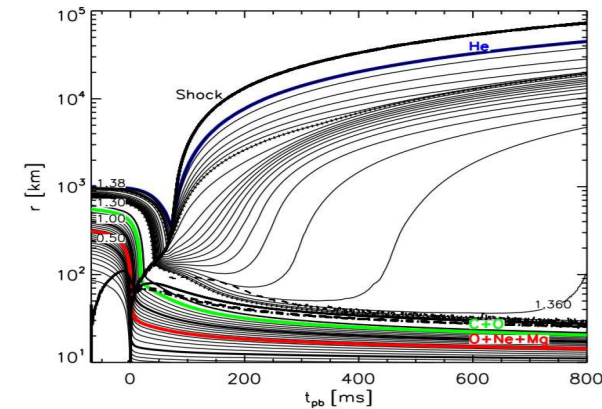
<http://www.franciscokitaura.wix.com/home>

I am Spanish-Japanese (born in Madrid), but raised in the German culture (school+University).

PhD at the Max Planck Institute MPA Munich.

Before coming to AIP I was Marie-Curie fellow in Italy (Trieste, Pisa) and PostDoc at the Excellence Cluster/MPE (Munich).

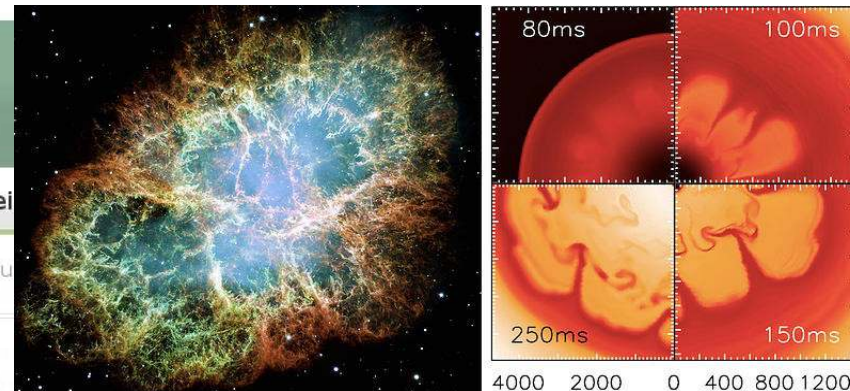
I started working in Astrophysics on **Supernova explosions**



SPIEGEL ONLINE WISSENSCHAFT

Politik | Wirtschaft | Panorama | Sport | Kultur | Netzwelt | Wissenschaft | Gesundheit

Nachrichten > Wissenschaft > Weltall > Supernova-Explosion: Forscher simulieren Krebsnebel-Geburt

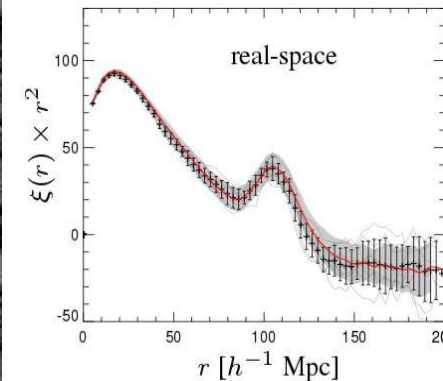
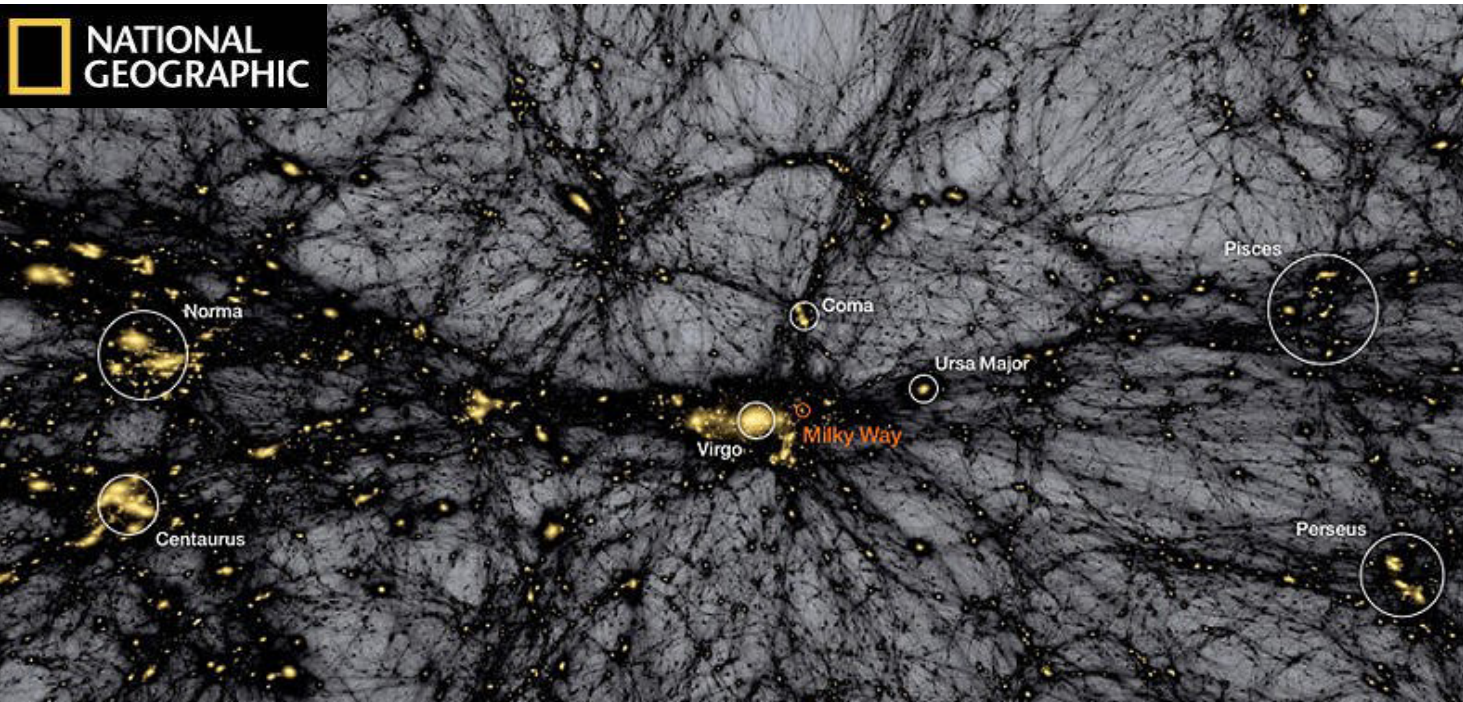
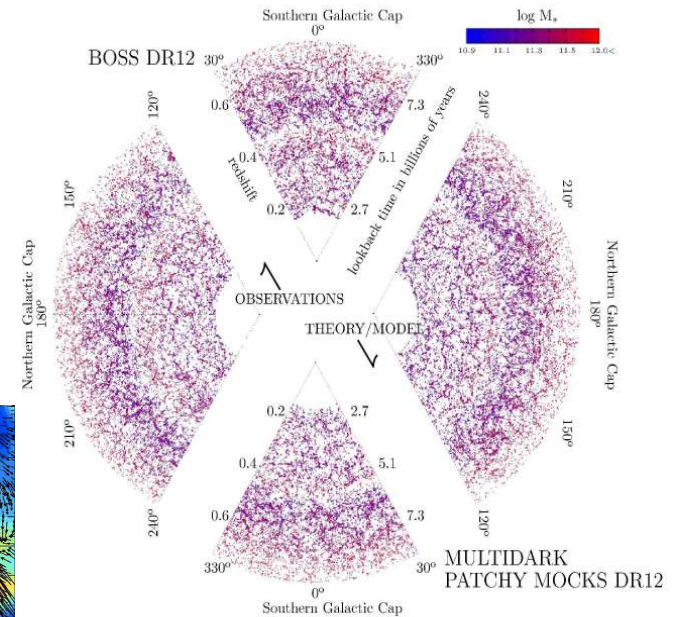
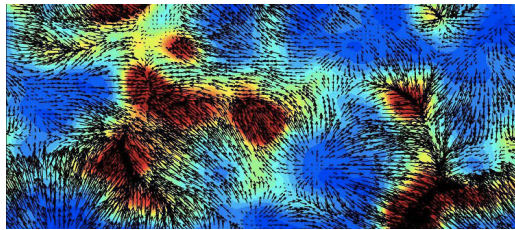


Supernova-Explosion: Forscher simulieren Krebsnebel-Geburt

Die Energie von Neutrinos zersprengte vor rund tausend Jahren einen Neutronenstern - dabei entstand der Krebsnebel. Max-Planck-Forscher konnten diese gigantische Supernova-Explosion jetzt erstmals am Computer simulieren und so eine alte Theorie beweisen.

Cosmology: Theory+Data Analysis

- PI cosmology science case 4MOST
- Co-lead BAO reconstruction WG EUCLID
- Co-lead Data Analysis WG J-PAS
- Working for BOSS on galaxy mock production and clustering analysis.
- Work on the Local Universe

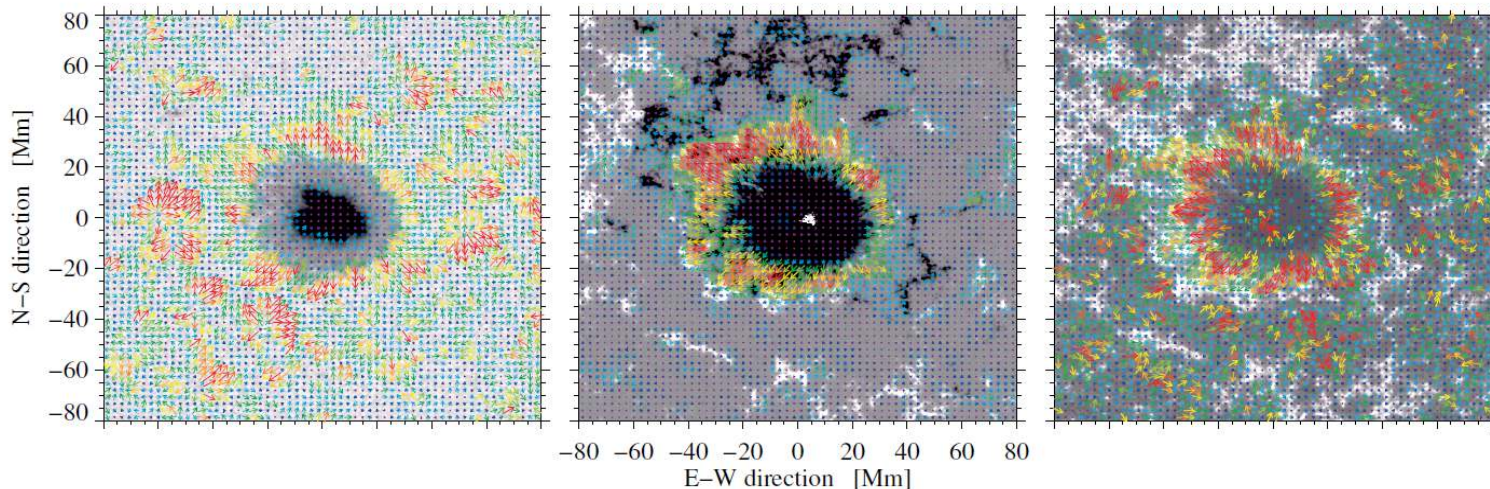


Philipp Kummerow

Solar Optical

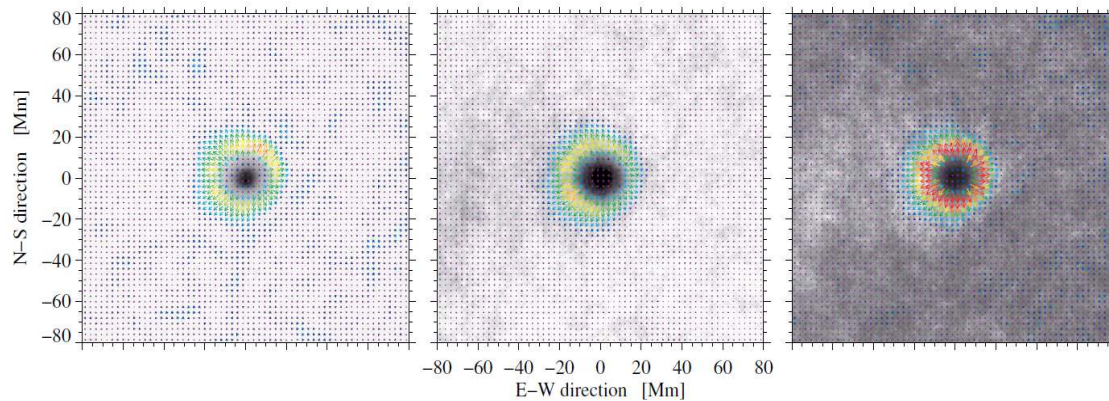
Horizontal proper motions in and around “theoretican’s sunspots”

- Philipp Kummerow
- Master’s student at the university of Potsdam
- In the Optical Solar Physics Group

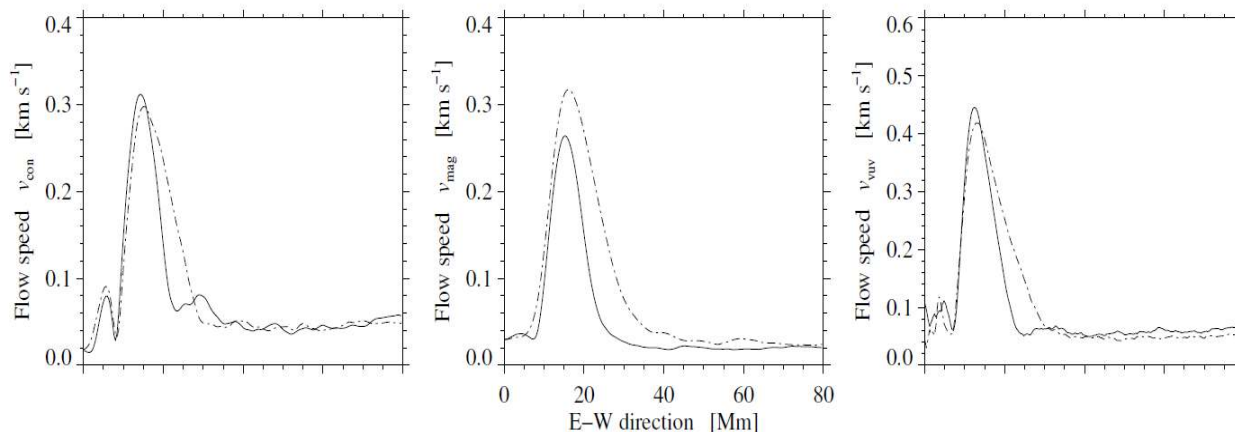


- Continuum intensity I , magnetic flux density B and EUV $\lambda 160$ nm for 26 axis-symmetric sunspots
- rainbow-colored vectors represent magnitude and direction of the horizontal proper motions

East-west asymmetry of sunspot moats



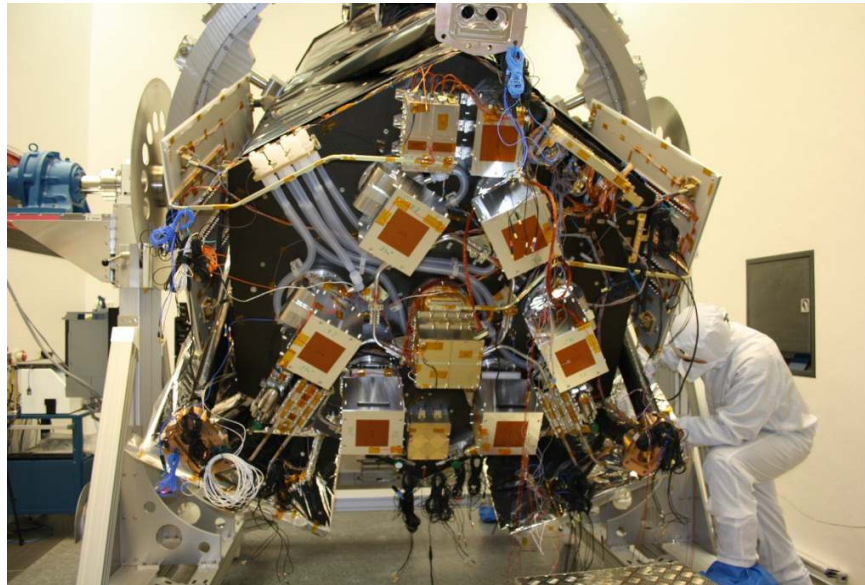
- Average sunspot
- moats are clearly asymmetrical in the east (dash-dotted)-west direction



Georg Lamer

Galaxies

Data analysis software for X-ray missions

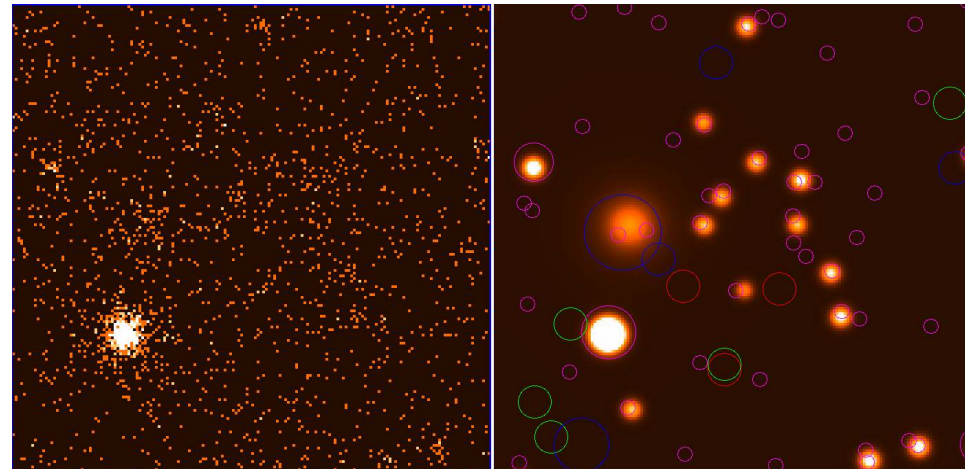


eROSITA focal instrumentation

**eROSITA simulation and
source detection**



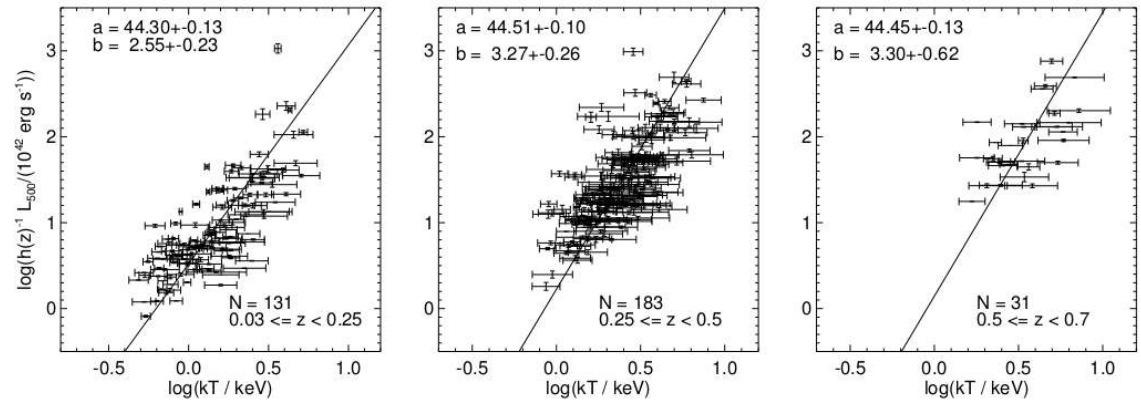
XMM-Newton



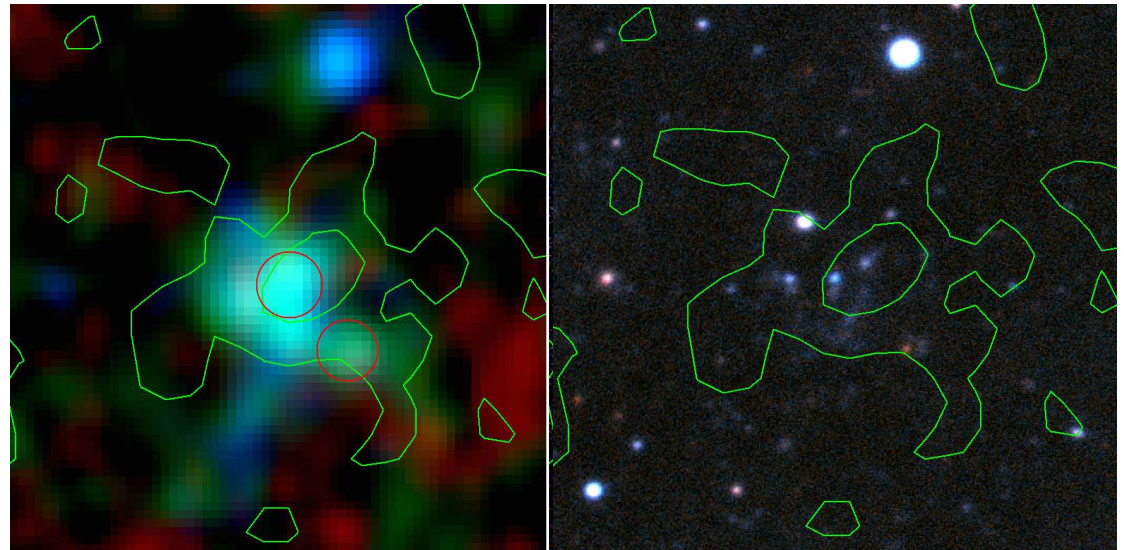
X-ray surveys for galaxy clusters



Massive cluster at $z = 0.99$
XMM/LBT



X-ray luminosity-temperature relation for 2XMM-SDSS clusters
(Takey et al. 2013)



X-ray contour on
WISE IR image

LBT r-z image

Umberto Maio

Cosmology

The birth of the first stars and galaxies

Umberto Maio

umaio@aip.de

cosmology group



Leibniz-Institut für
Astrophysik Potsdam

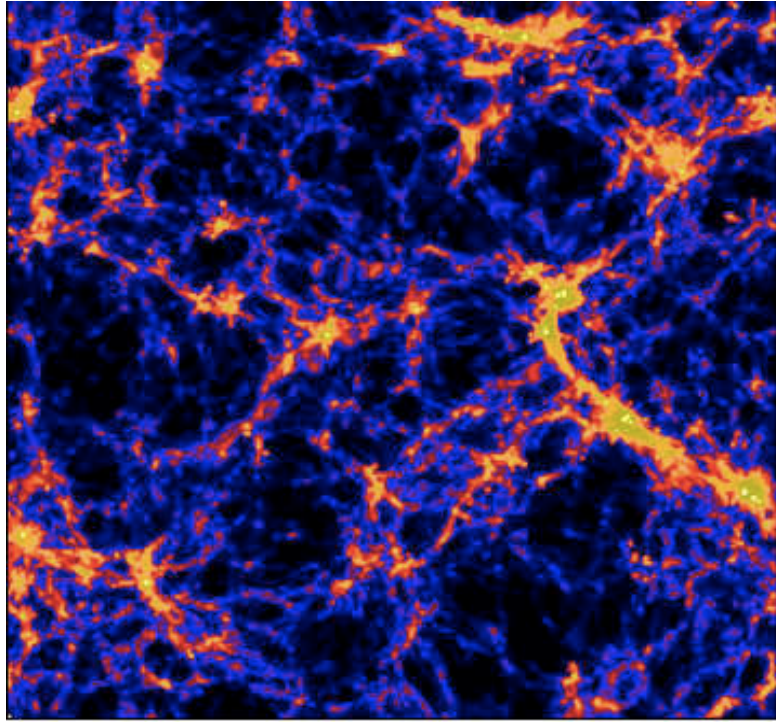
In collaboration with: K. Dolag, B. Ciardi,
M. Petkova, L. Tornatore, S. Khochfar, J.
Johnson, R. Salvaterra, N. Yoshida, V.
Müller, V. Biffi, L. Koopmans, M. Viel, ...

Goals

- Early Structure evolution
- Formation epoch
- Astrochemistry – molecules & Z
- Role of popIII and popII
- Comparisons to observables

Relevance

Upcoming international missions:
SKA, JWST, E-ELT



Simulations

N-body + hydro + cooling & SF

Non-equilibrium molecular chemistry
(e^- , H, H^+ , H^- , He, He^+ , He^{++} , H_2 , H_2^+ , D, D^+ , HD, HeH^+)

Metal spreading for He, C, N, O, Si, S, Mg, Ca, Ne, Fe, etc.

Different cosmologies (Λ CDM, non-Gaussian, warm dark matter, ...)

Observables:

- Luminosity functions
- SFRD and sSFR
- Chemical features

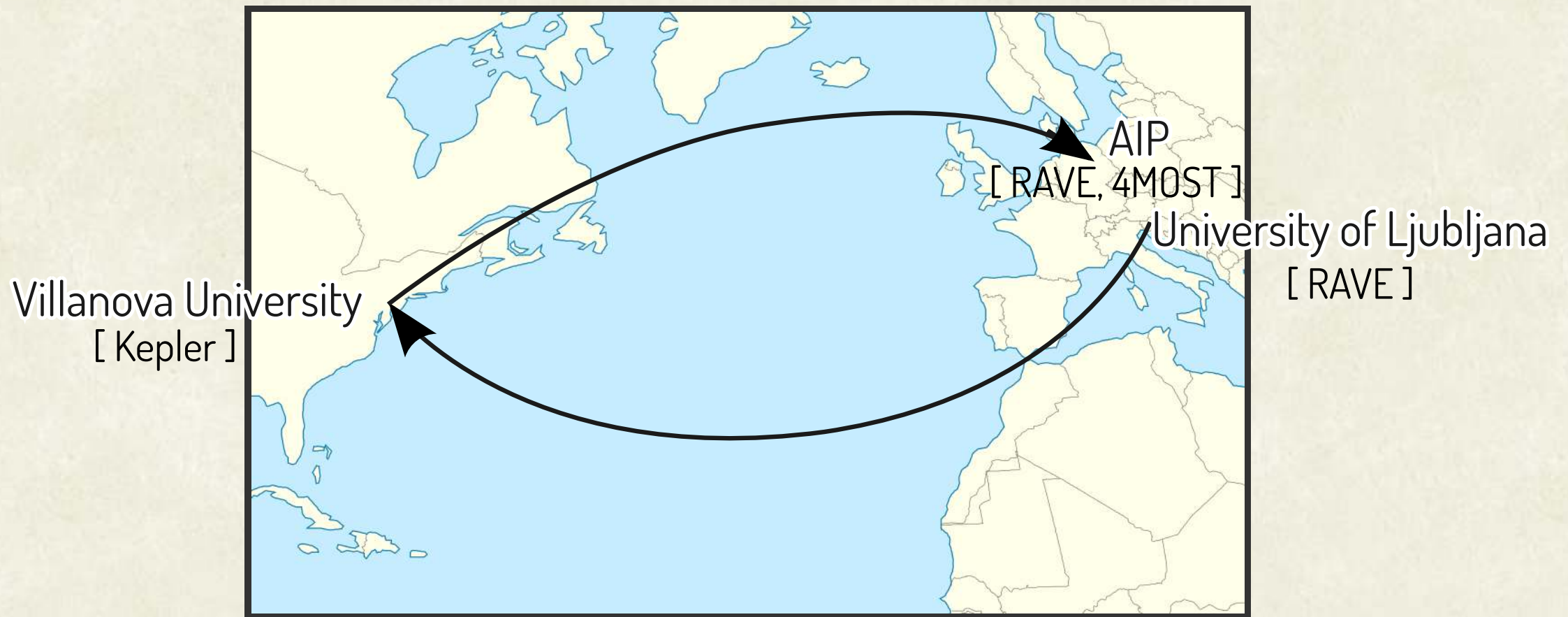
Theoretical Predictions:

- Baryon properties in early objects
- Implications for GRB and DLA hosts
- Effects of different cosmologies

Gal Matijevic

Milky Way and Local Volume

Gal Matijević



application of

(probabilistic) data analysis &
machine learning techniques

to astronomical
datasets

RAVE Explorer

t-SNE projection of RAVE spectra



- ☒ None
- ☐ Temperature
- ☐ Gravity
- ☐ Metallicity
- ☐ S/N
- ☐ Density
- ☐ Radial vel.
- ☐ Distance
- ☐ LLE Flag
- ☐ EWrt

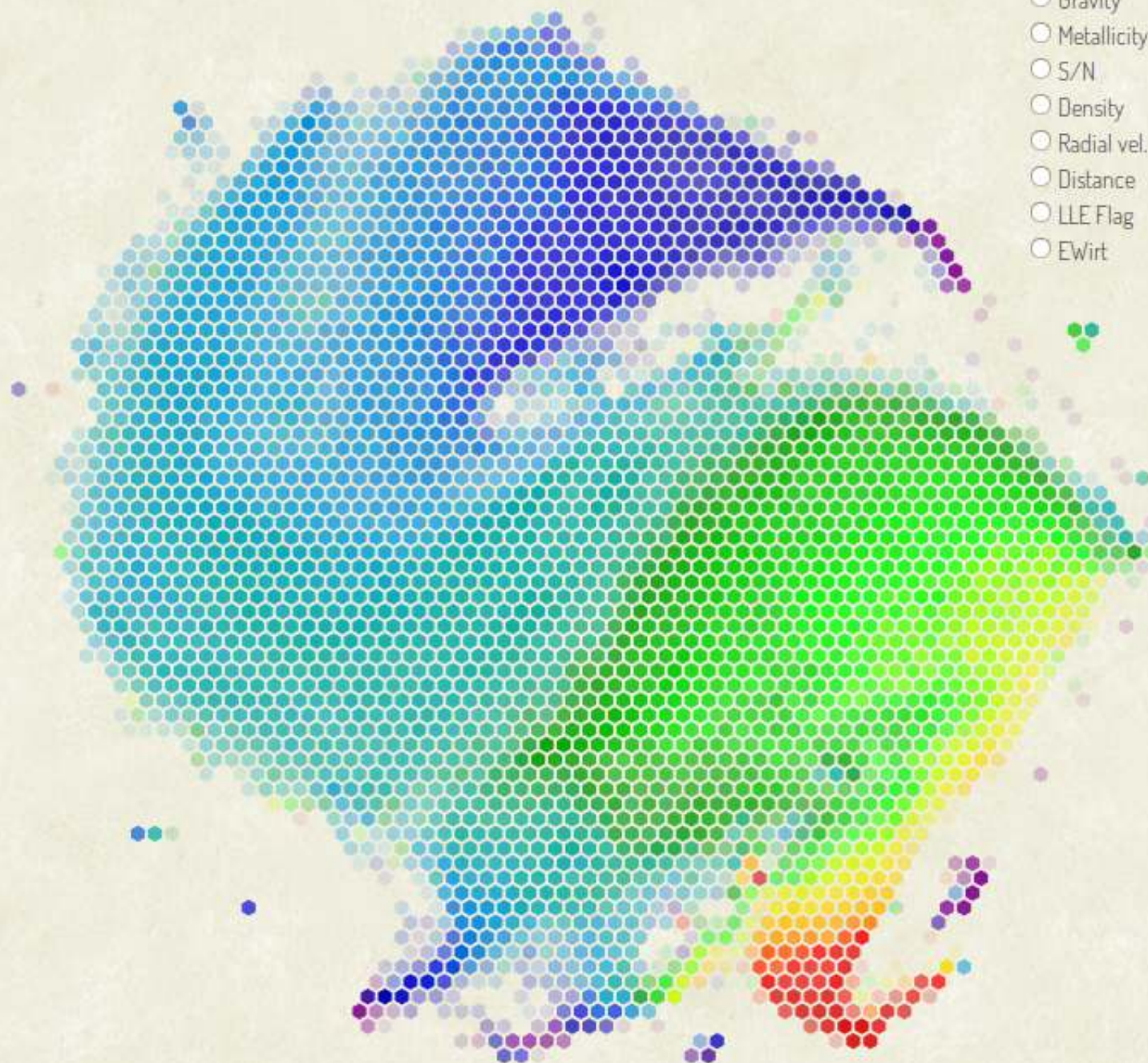


RAVE Explorer

t-SNE projection of RAVE spectra



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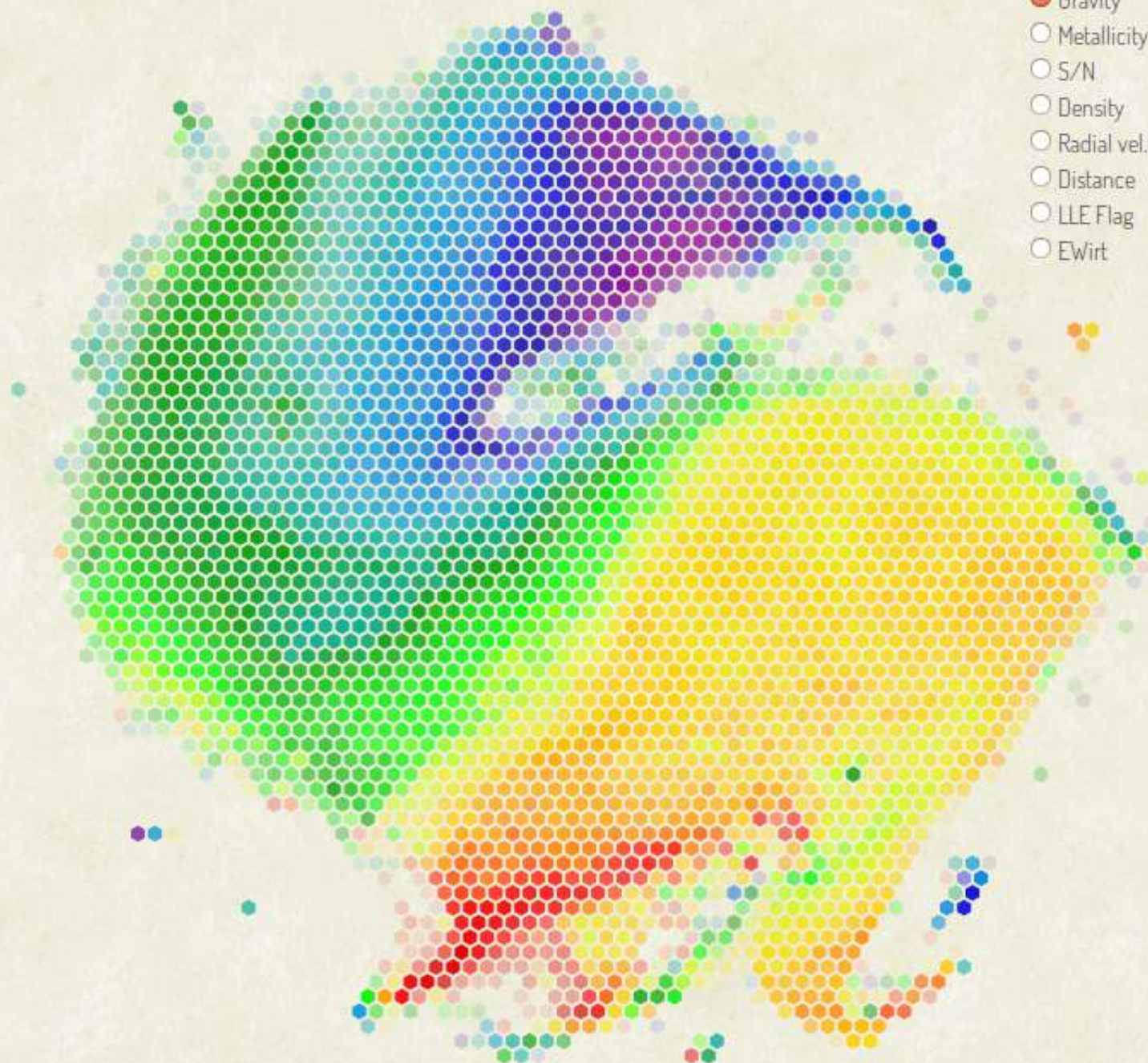


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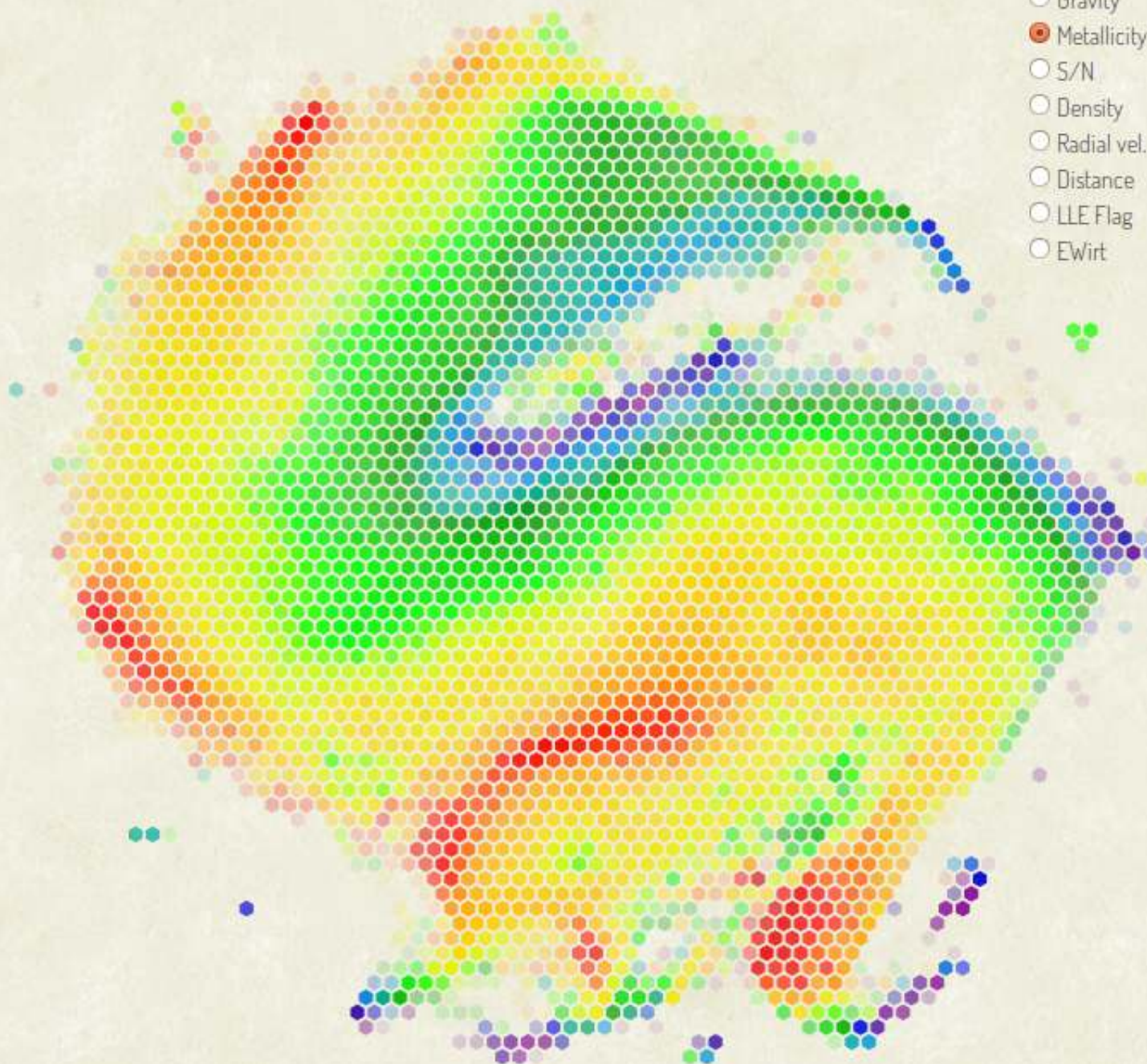


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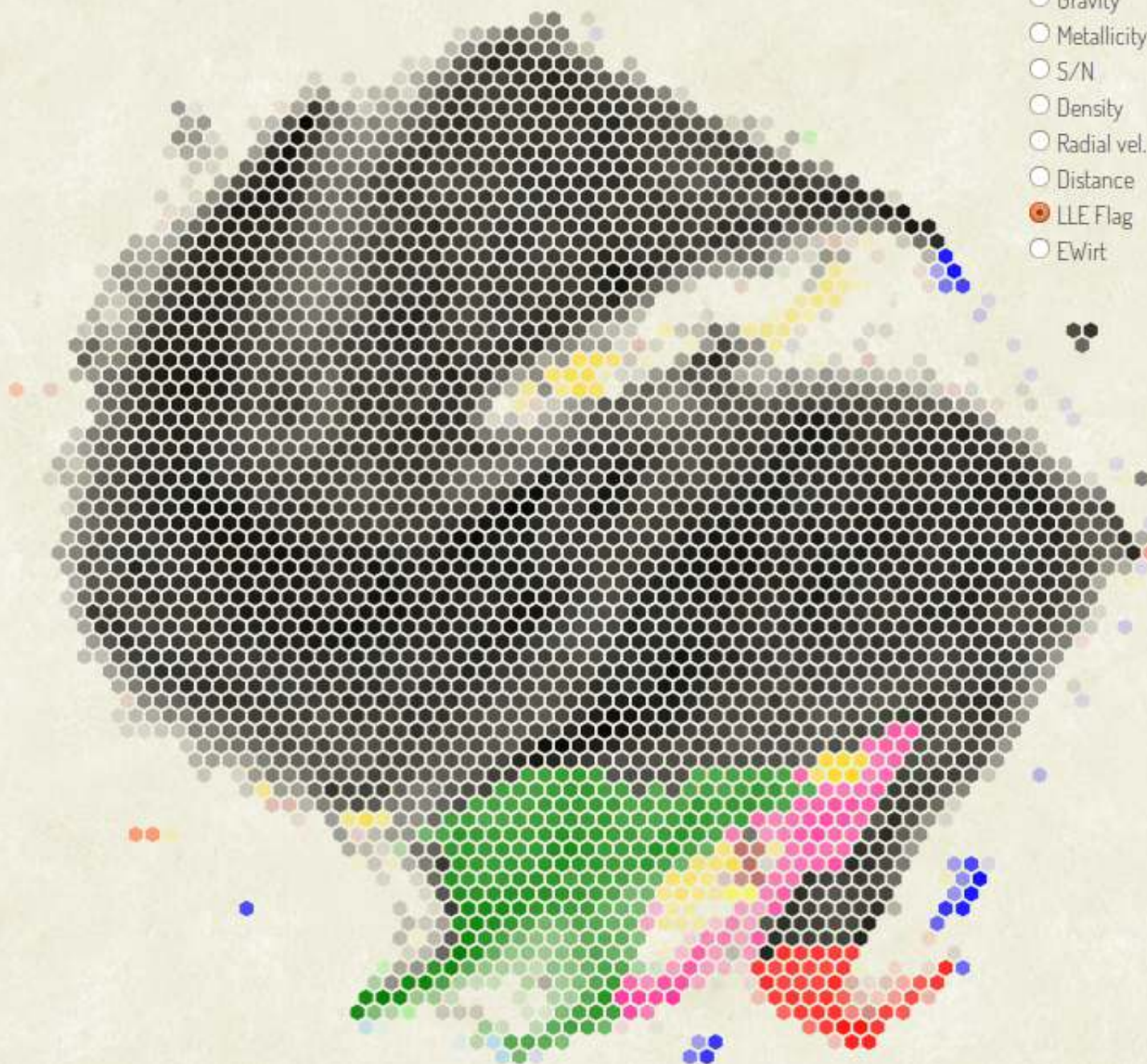


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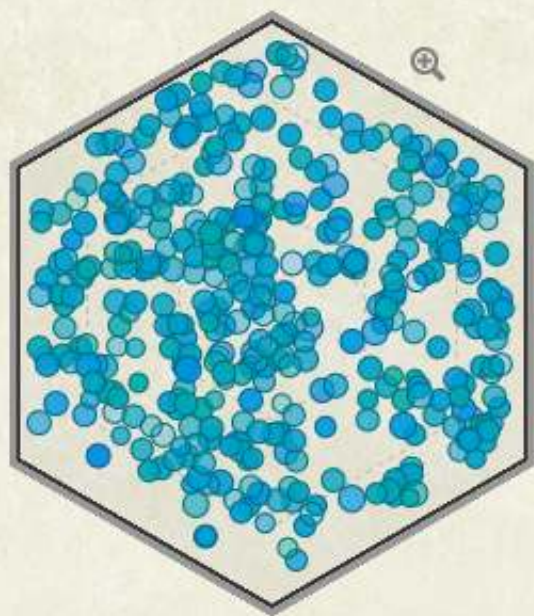


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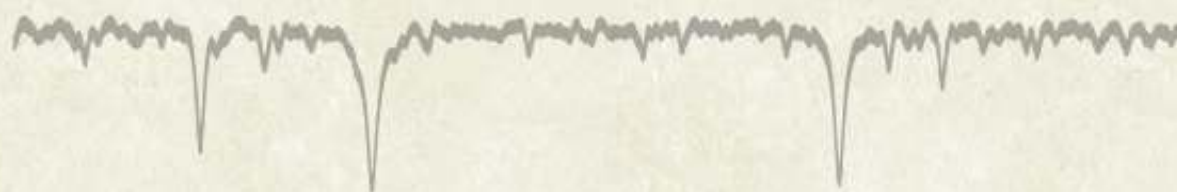
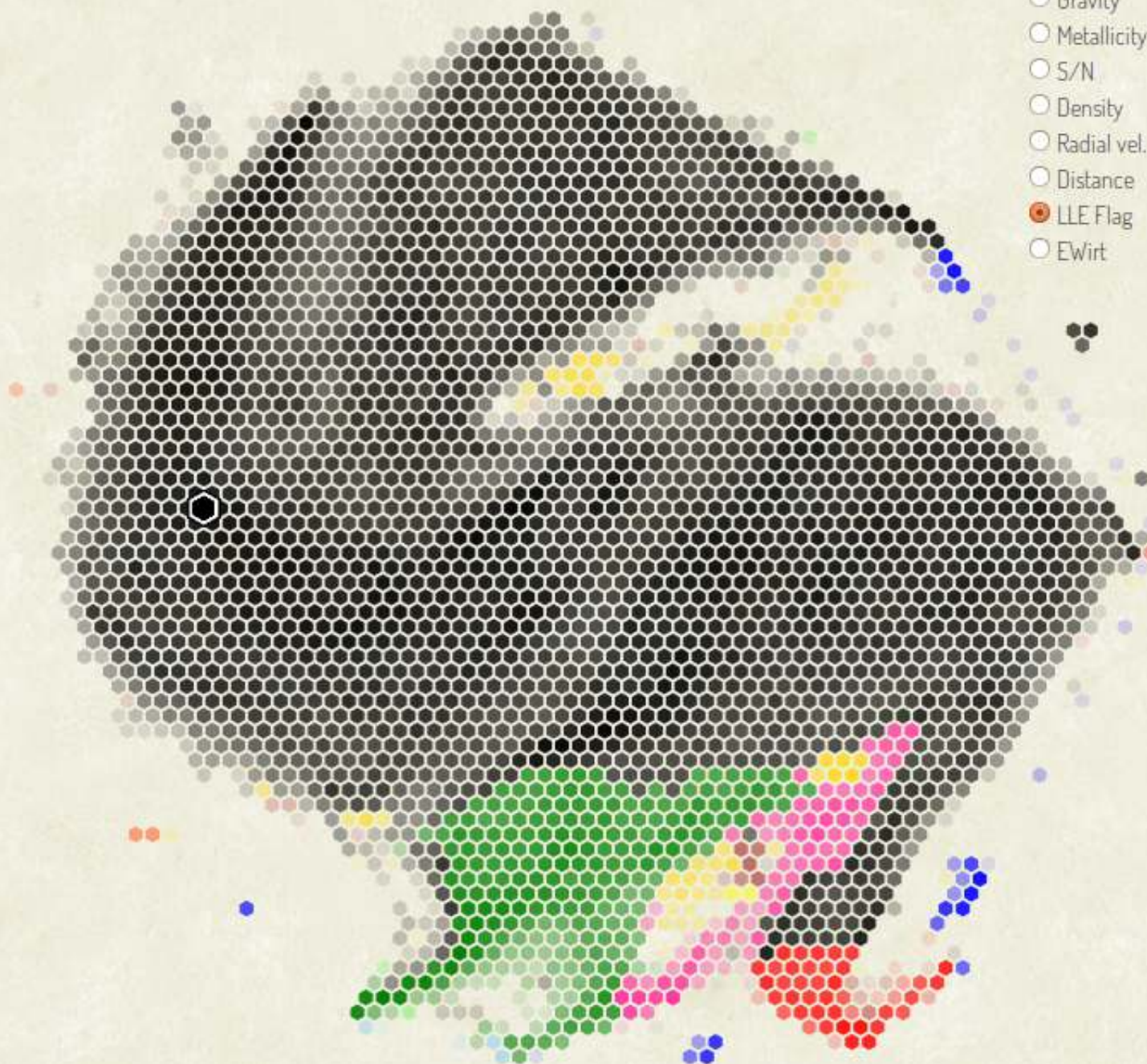


RAVE Explorer

t-SNE projection of RAVE spectra

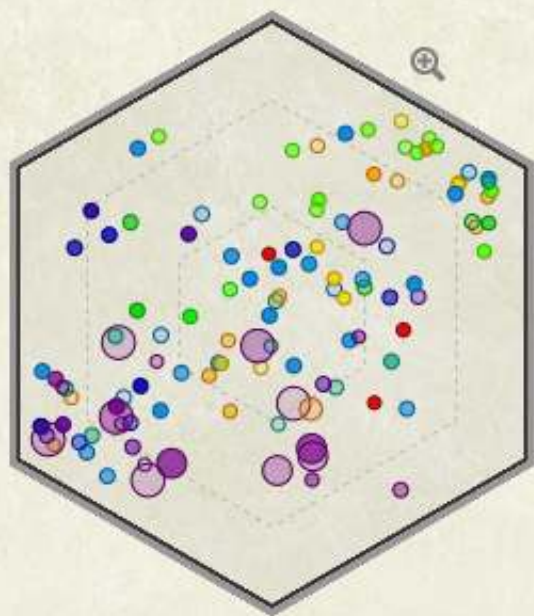


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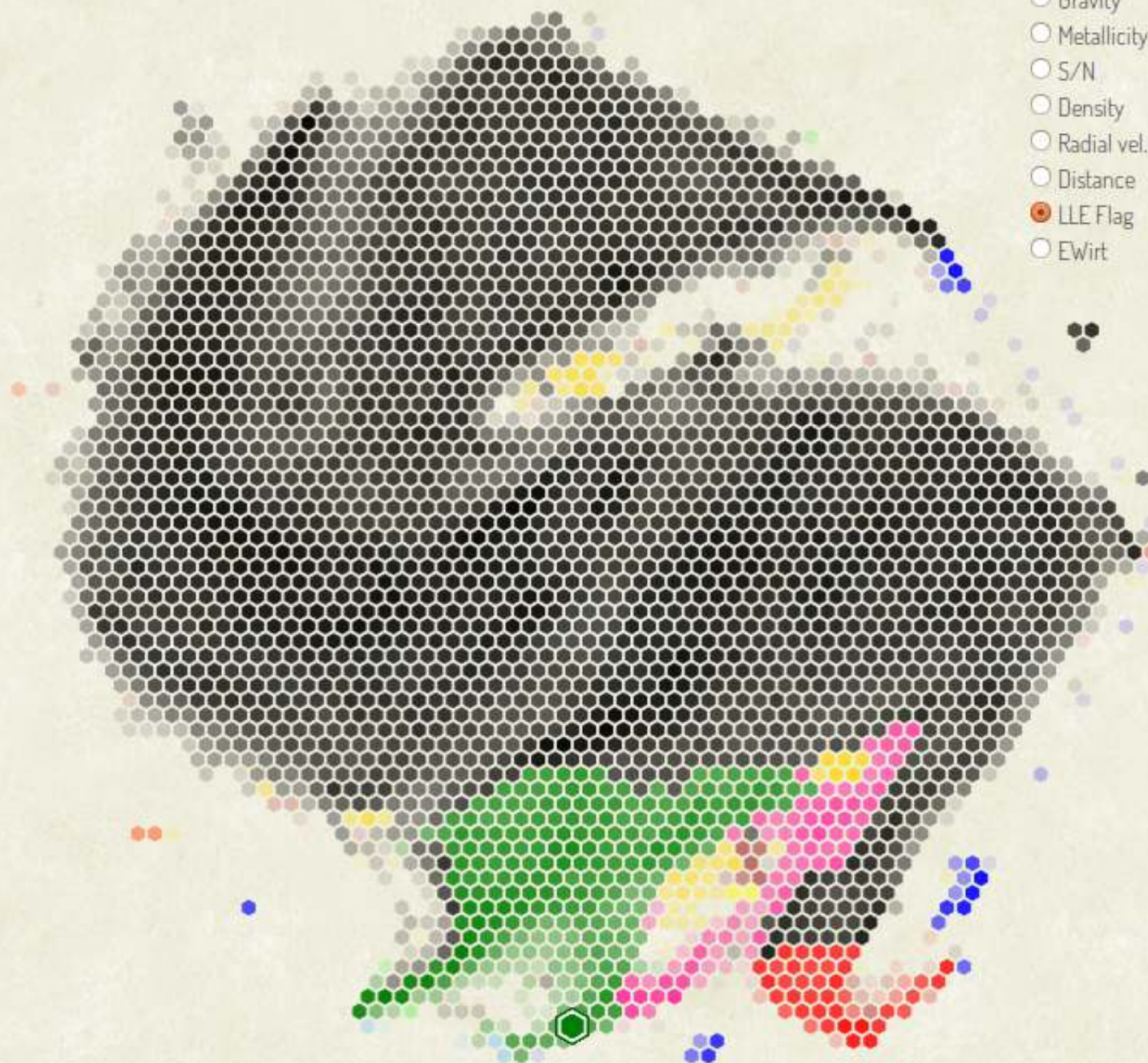


RAVE Explorer

t-SNE projection of RAVE spectra

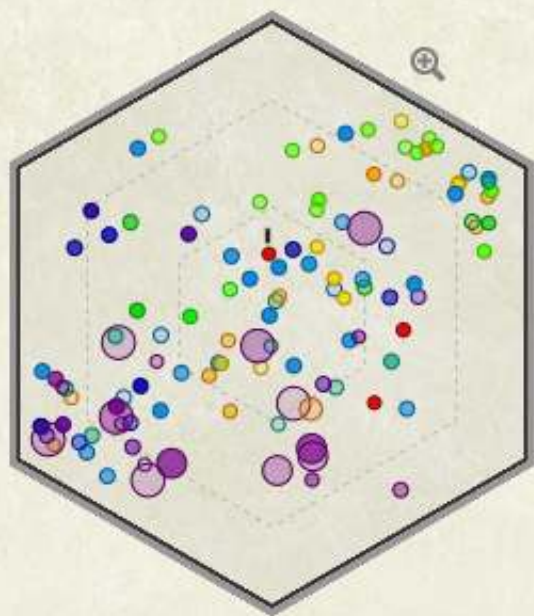


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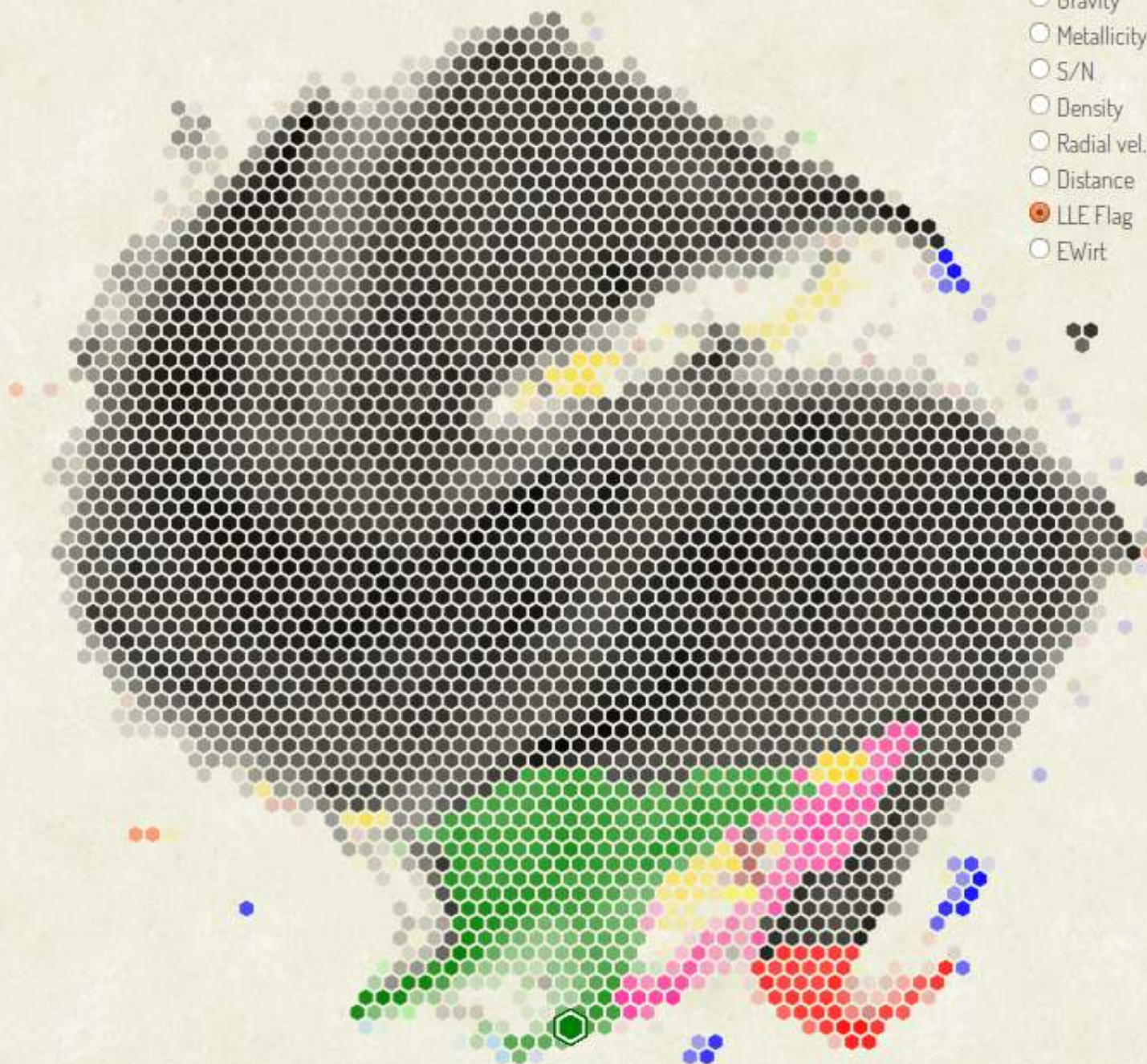


RAVE Explorer

t-SNE projection of RAVE spectra



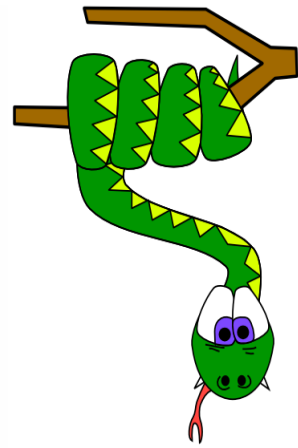
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- EWirt



Alexey Mints Galaxies



Who I am: Alexey Mints



What I do:
The Integrated
cluster finder

ARCHES cross-matching tool



SDSS + UKIDSS + AllWISE catalogs



Redshift-color relation



Cluster candidates



3XMMe + Cluster finder

[Full list](#)

XMM Source ID: 129421

ICF iid: 24773

Detections data:

Catalog	Redshift	Redshift error	N spectra	Redshift (spectroscopic)	Lambda	BCG distance (Mpc)	Weighted radius (Mpc)
xmatch	0.2157	0.0302	9	0.21798	27.75	0.97	0.08

Cross matches:

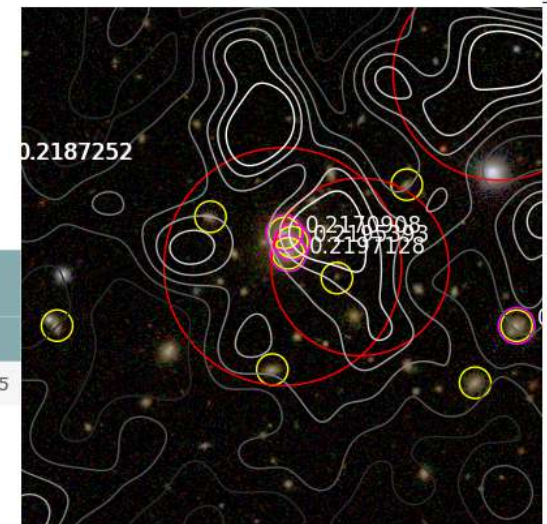
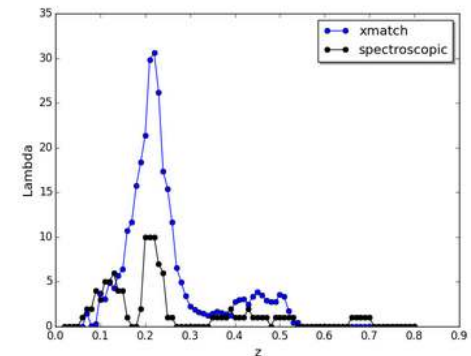
Catalog	ID in the catalog	Redshift	Separation(arcmin)
Takey2	8084	0.21	0.06
RedMapper	3514	0.22	0.17
MAXBCG	8495	0.23	1.57

3XMM detections

Detection ID	OBS ID	Flux (Band 8)				Detection likelihood (Band 8)				EP Extent		
		EP	PN	M1	M2	EP	PN	M1	M2	extent	error	ML
21251	0303562201	2.382e-13	2.055e-13	8.871e-13	2.015e-13	70.32	34.59	16.60	19.31	29.6653	3.39529	29.415

XSPEC fits

Redshift	OBS_ID	kT	kT (lower)	kT (upper)	L
0.21567	0303562201	2.497	-0.669	1.261	9.494e+42
0.21567	total	2.497	-0.669	1.261	9.494e+42



Martin Mitzkus
innoFSPEC

Martin Mitzkus – PhD student

Göttingen:

Bachelor & Master of Physics

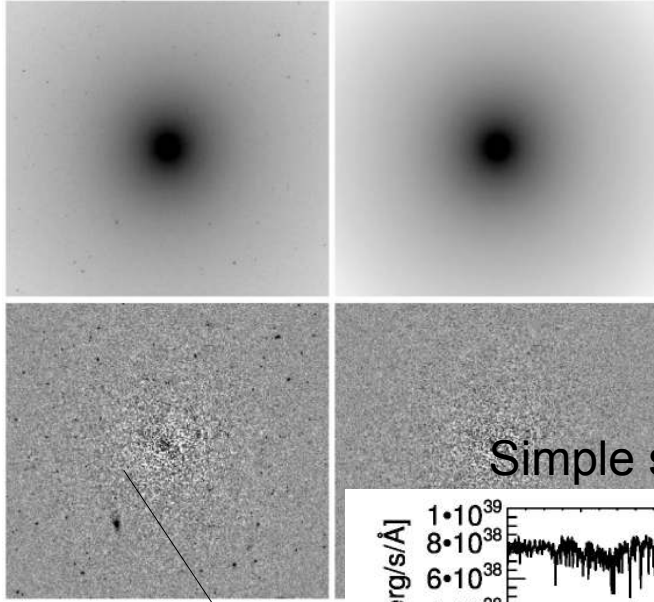


GEORG-AUGUST-UNIVERSITÄT
GÖTTINGEN



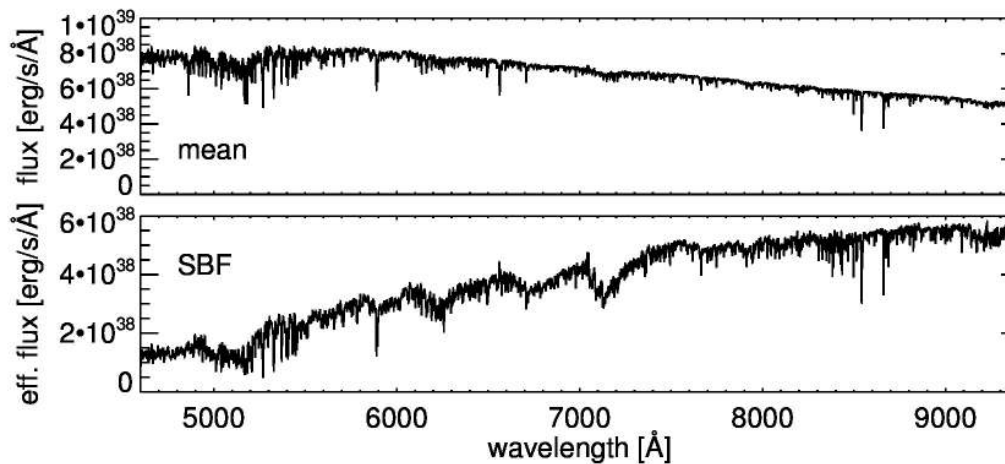
Spectroscopic SBF

Surface Brightness Fluctuations

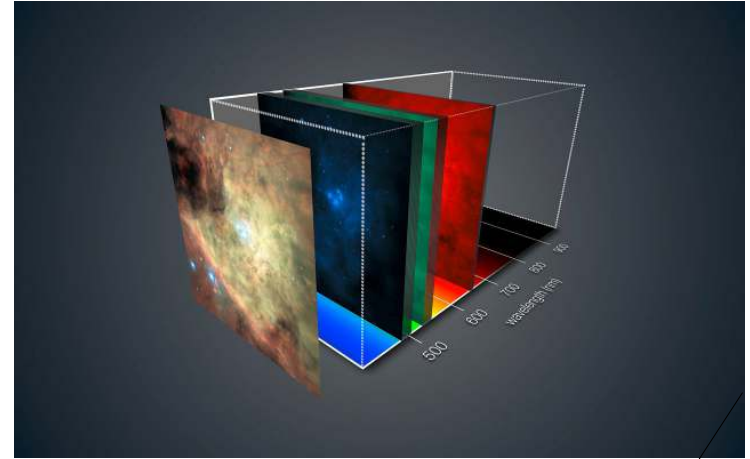


Cantiello et al. 2005

Simple stellar population (10 Gyr, $z=0.004$)



Integral field spectroscopy: MUSE



Credit: ESO

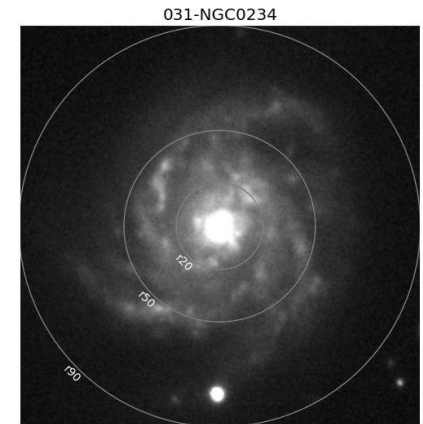
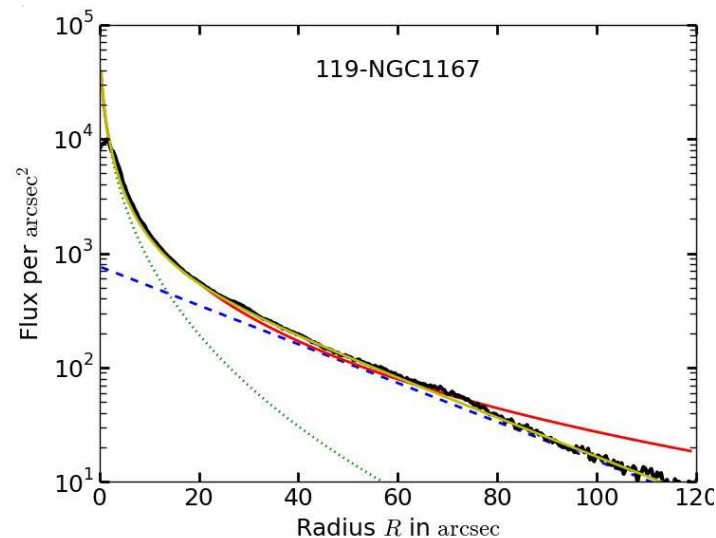
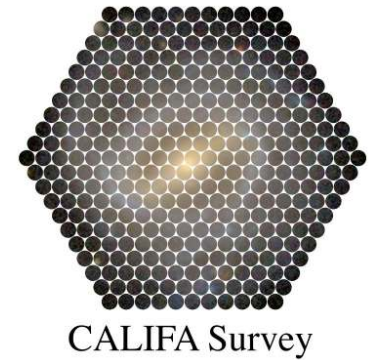
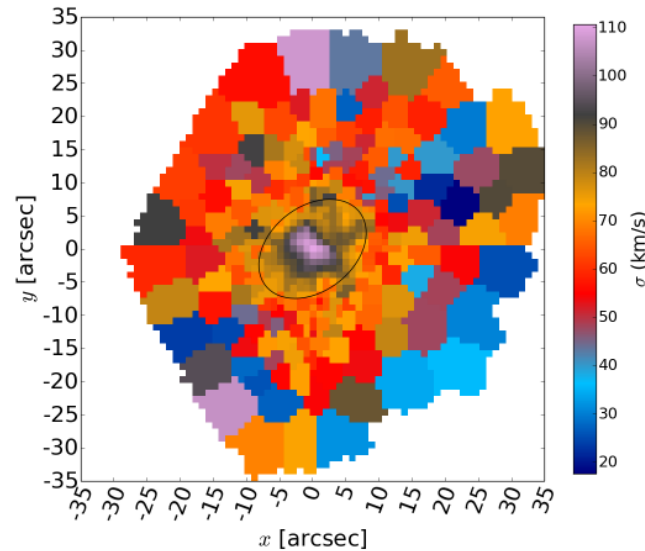
Justus Neumann

Galaxies

Bulges in CALIFA Galaxies

Diploma
Thesis

- Detection and classification of bulges
- Using visual, photometric and spectroscopic approaches



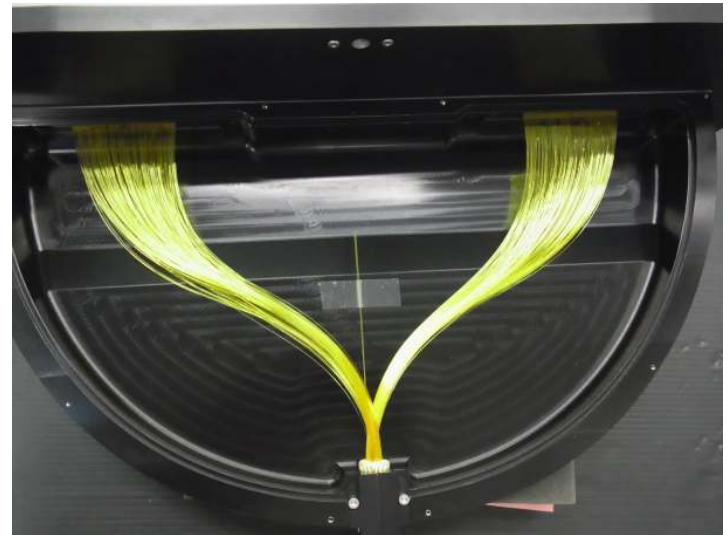
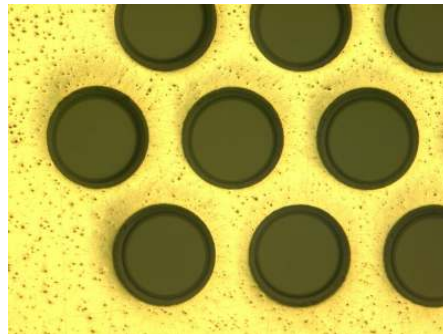
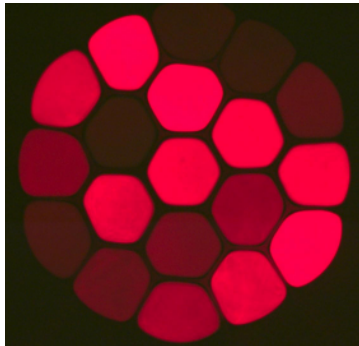
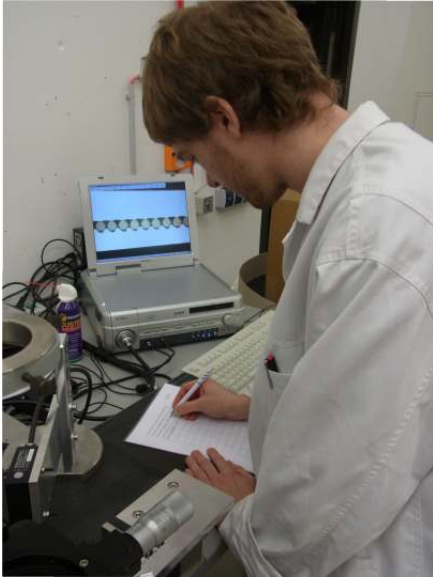
VIRUS Quality Control

Student
Assistant

+ ERAMUS-F & 4MOST

Quality control of
VIRUS fibre bundles:

- Polishing inspection
- Focal ratio degradation evaluation
- Relative throughput measurement



Alejandro Paredes

MHD

Something about me:

- **Career path**

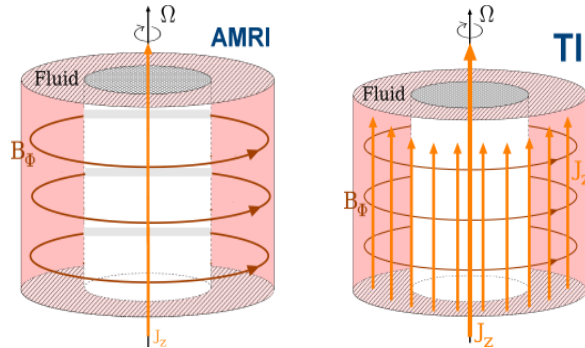
- **BSPHy.** **National University of Engineering (Peru 1998 -2004)**
- **Master of Physics** **Aix-Marseille Université** **(France 2007-2009)**
Theoretical physics
- **Phd. Fluid Mechanics** **Aix-Marseille Université** **(France 2009- 2013)**
Numerical simulations for edge plasma in tokamaks
- **Postdoc MHD group** **AIP - MHD group** **(Germany 2013-)**
Numerical simulations for MRI

- **Interests:**

Computational Fluid Dynamics, High Performance Computing

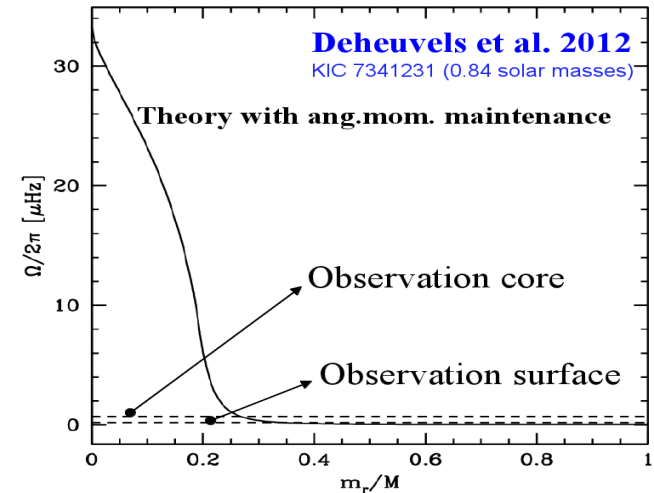
My topic

- MHD Numerical Simulations
- MRI



- FSEM code :
MHD in Taylor Couette set-up

Why?



Candidate to explain disagreement

My contribution

Implement a solver in FSEM code : MHD + $Pm \ll 1$

This means to solve : $\nabla^2 \vec{b} = \vec{F}$, $\nabla \cdot \vec{b} + \text{B.C.}$

Daniel Phillips

Milky Way and Local Volume

What have I been doing Professionally?

- Educational Background: Bachelor of Science & Master of Civil Engineering
- 13 yrs construction & contracting, NE USA



- 4 yrs with Cerro Tololo Inter-American Observatory in La Serena, Chile; Facilities Group



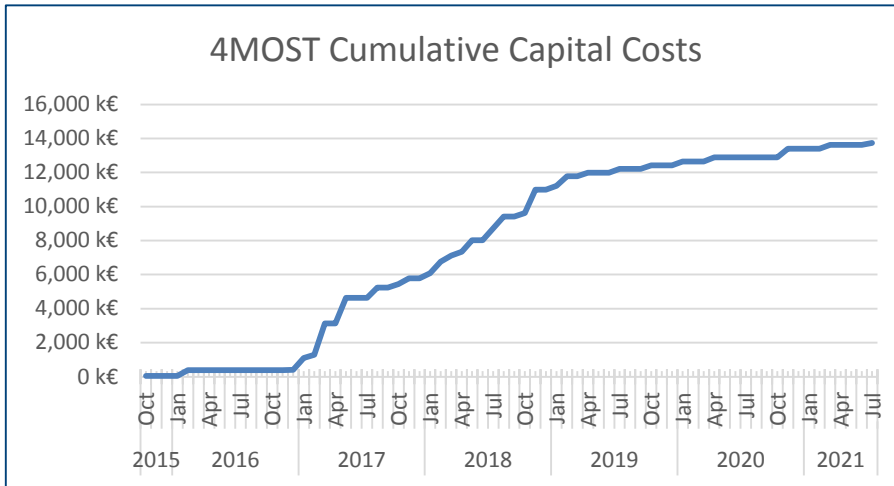
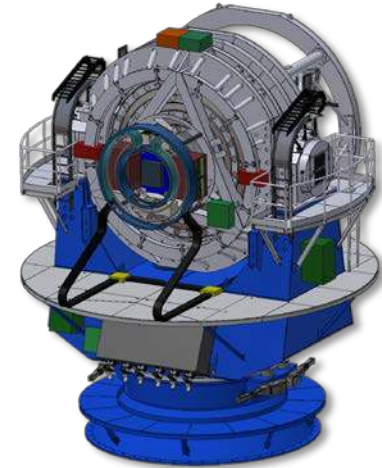
- 1 year with LSST; Telescope Site Leveling Project

- Now: Project Controller for 4MOST
 - Range of Project Management Activities in the 4MOST Project Office
 - Scheduling
 - Budget Control
 - Developing Project Management documents and tools





- 4MOST has completed the Conceptual Design Phase.
- Preliminary Design Phase Kick-off: Just done!
- AIP contributes >40% of the staff effort and ~25% of the HW by Capital Cost, + Science
- Project Control: Support On-time and On-Budget Completion of 4MOST by interacting closely with PM and consortium partners.



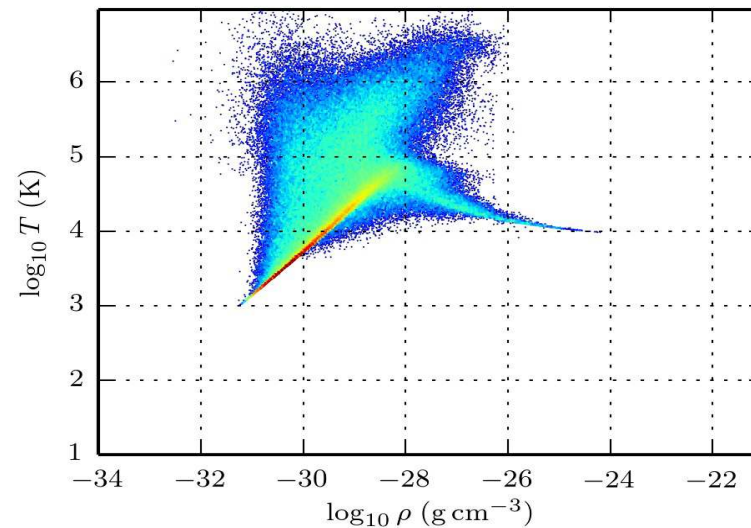
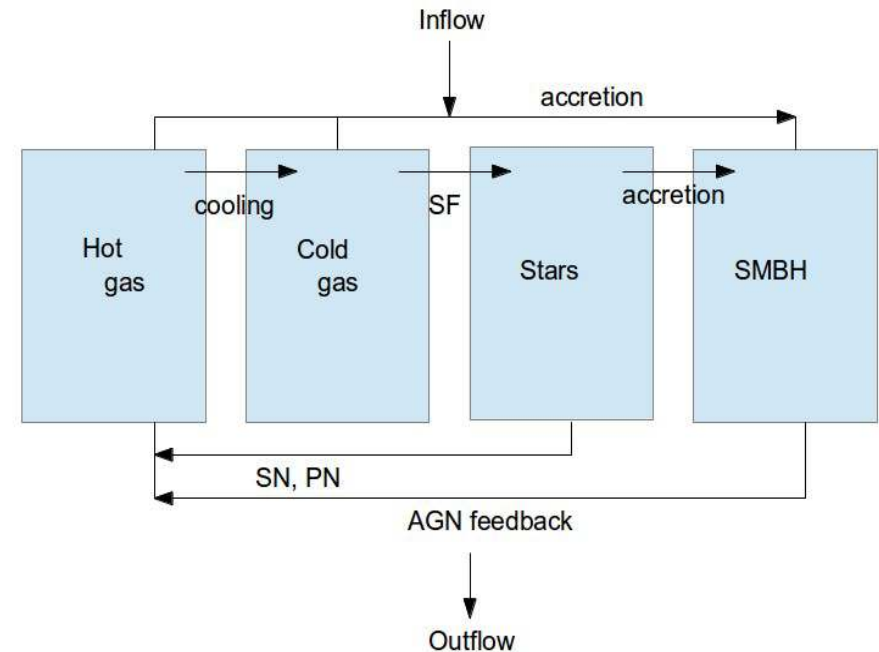
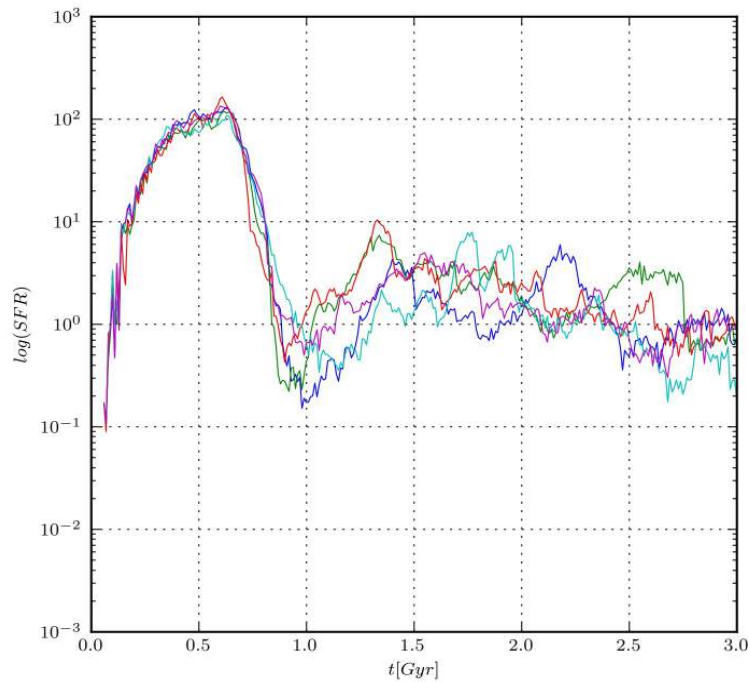
Project Phase	Staff FTE
Preliminary Design	41.4
Final Design	57.0
MAIV	117.3
Transport	2.7
Integration & Commissioning	21.6
Total:	240

Pierre-Antoine Poulhazan

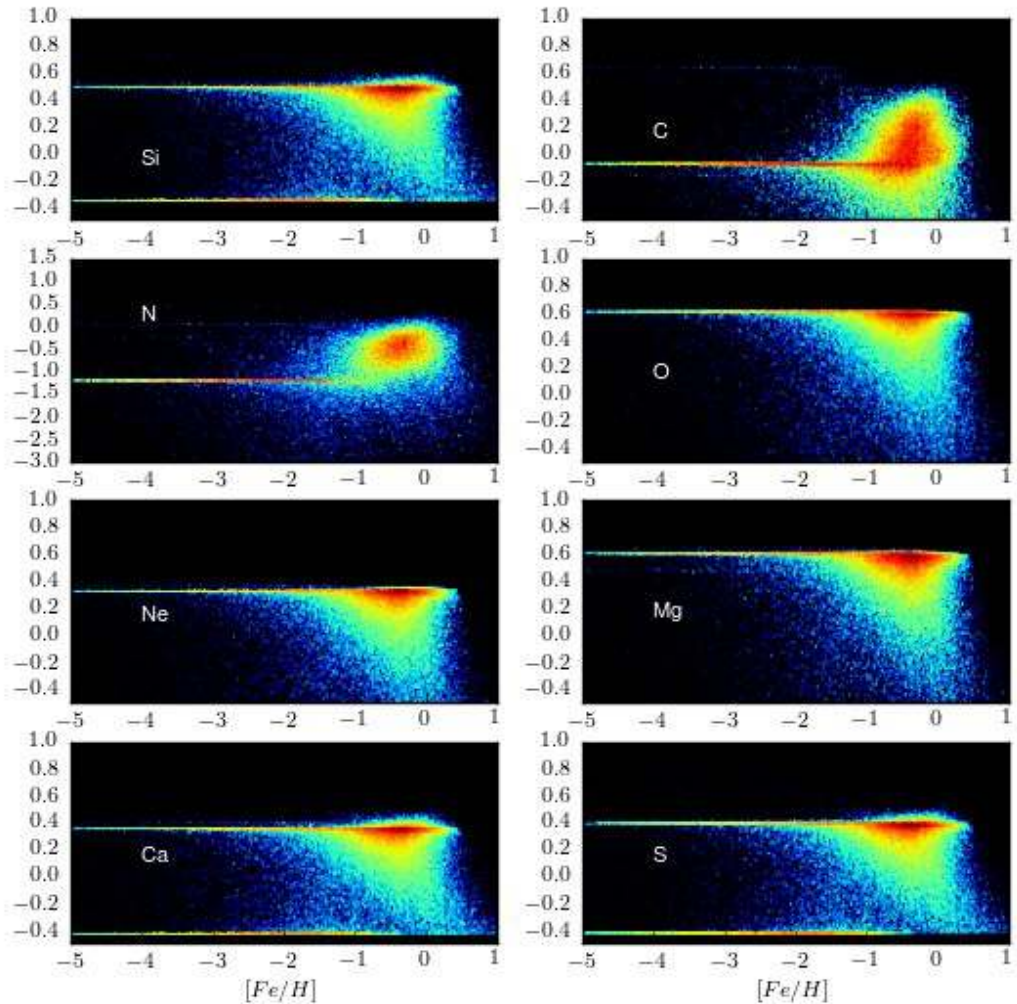
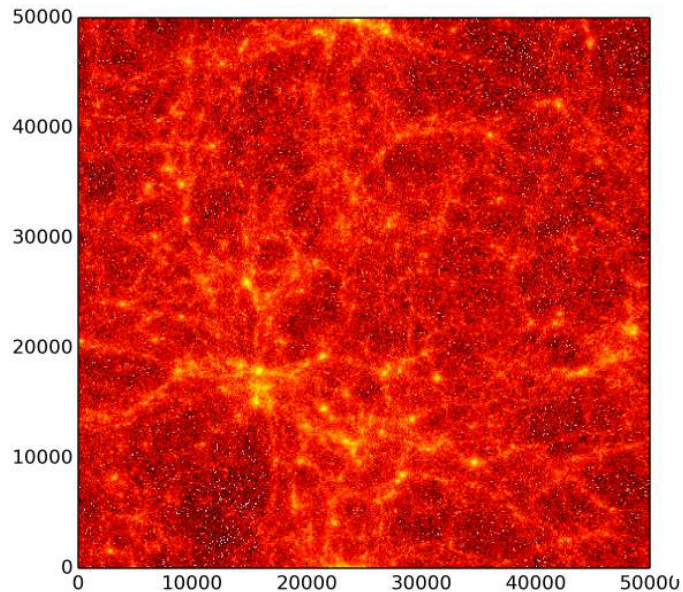
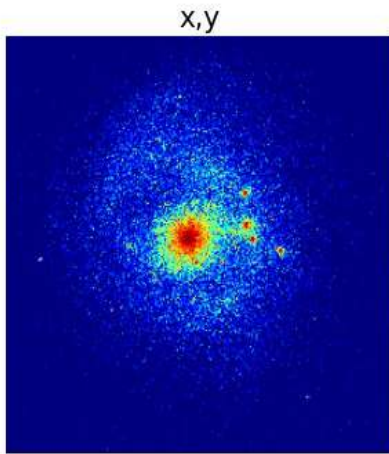
Milky Way and Local Volume

Chemical Enrichment in cosmological SPH simulations

- Initial Mass Function
- Life-times of progenitor stars
- Rates of SNIa, SNIi and AGB events
- Chemical Yields
- Cooling



Chemical Enrichment in cosmological SPH simulations



Andreas Rabitz

Galaxies



Background

- AIP history: student worker & Diploma thesis
- PhD student (galaxy clusters) – Axel Schwope
- observer (opt./near-IR imaging & spec.)

Recreation

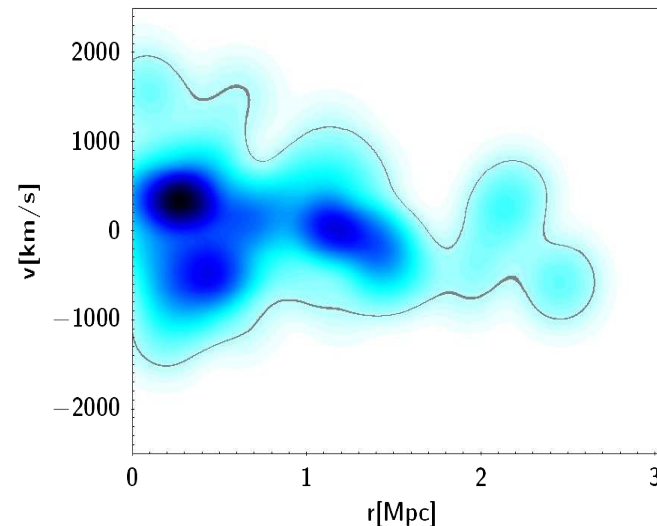
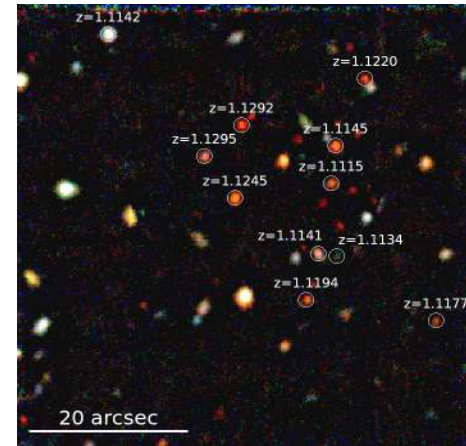
- biking, ca. $(1.2_{-3}^{+2}) \times 10^{17} \text{ Å} \cdot \text{yr}^{-1}$
- swimming

current projects

- Cluster survey (LBT)
(G. Lamer, A. Schwobe)
→ X-ray, opt.follow-up, spec.
→ IR-selection (PhD)
- MUSE-SV
(G. Lamer, A. Schwobe)
one shot – one cluster
- S1101 – detailed cluster study
(Y. Zhang)

Knowledge

- scripting: bash, STilts
- data reduction: imaging & spectroscopy
→ „Bertin-tools“, Theli, MIDAS (mosp, O2K), EZ, ...



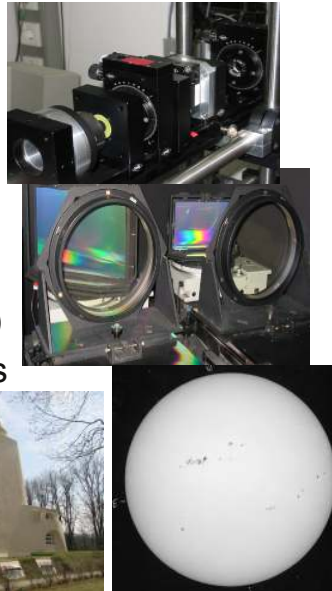
Jürgen Rendtel

Solar Optical

Solar Optical

Spectropolarimetric Lab Einsteinurm, Potsdam

- instrument setup and testing
- student lab.
(sunspot magnetic fields)
- digitization of solar plates

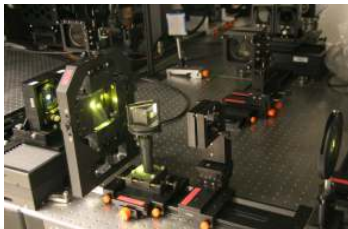


Telescopes (VTT & GREGOR)

and post focus instruments, Izaña, Tenerife



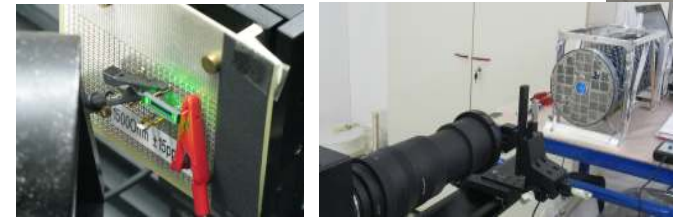
Polaris. Calibr. Unit
GFPI



Solar Radio

STIX team member

- optical tests for aspect system
- scheduling



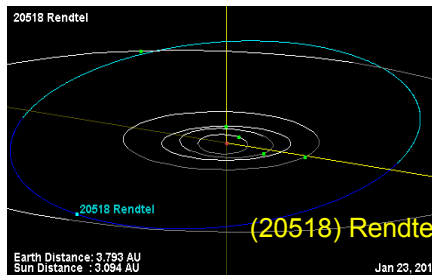
Other activities:

Enquiries from public & media

- astronomy in general
- transient phenomena

Internat. Meteor Org. (IMO)

Vice President

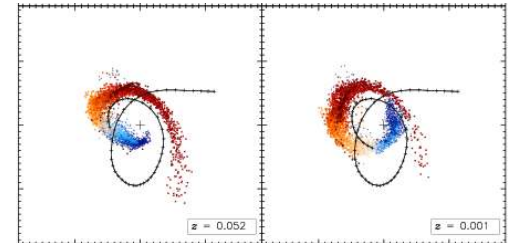


Kristin Riebe

e-Science

About me

- Studied physics in Heidelberg
- Diploma thesis at Astronomisches Rechen-Institut Heidelberg
- Short HPC-project at University of Edinburgh, EPCC
- PhD at AIP, cosmology group
- Since then: E-Science



Private

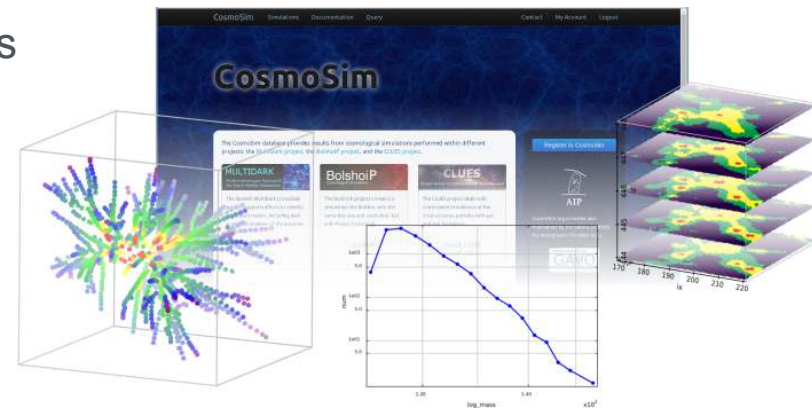
- Two children
- Computer graphics, illustrations



Work

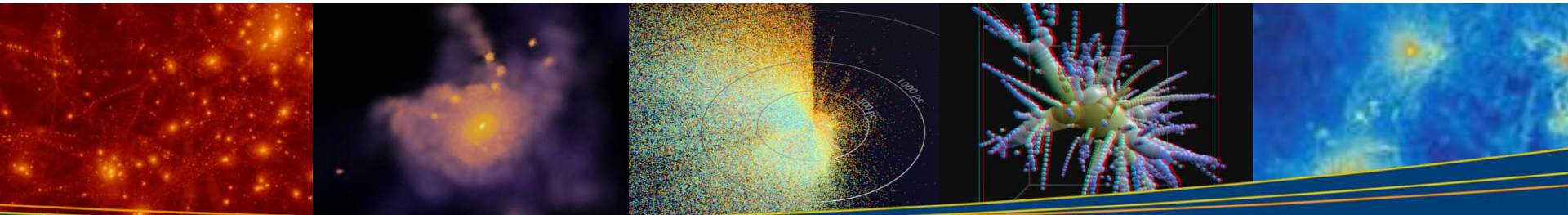
GAVO project

- German Astrophysical Virtual Observatory
 - Data publication
 - Databases, web services, standard definitions
- CosmoSim
 - database for cosmological simulations



Visualisation

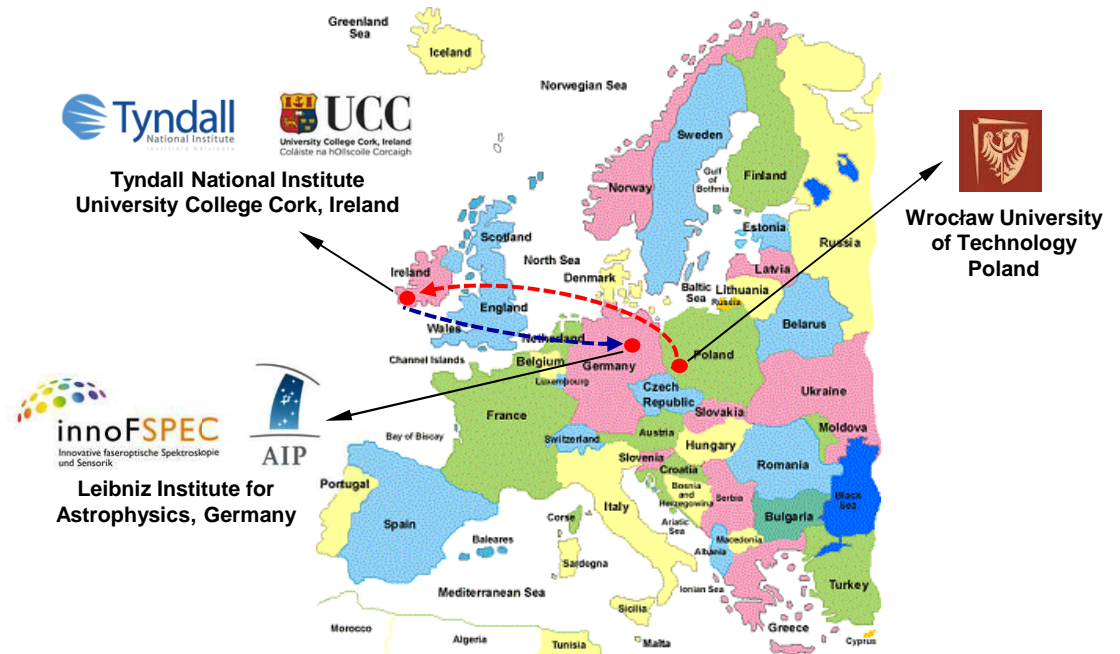
- cosmological simulations (PMViewer)
- RAVE stars, 3D demos (Blender)



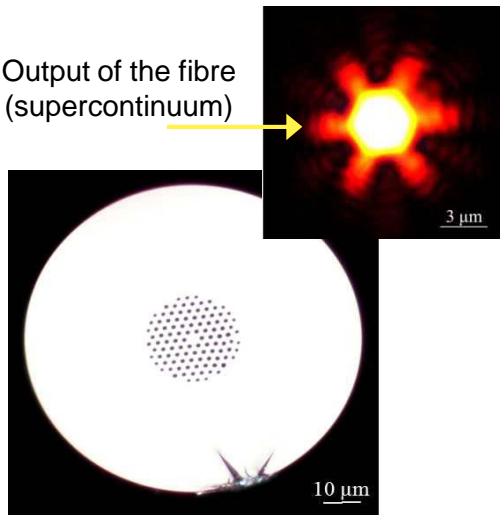
Monika Rutowska
innoFSPEC


```

/?/ who am I
monika
/?/ show tools
2003 M.Eng in Applied Physics Wrocław, Poland
2012 PhD in Physics Cork, Ireland
/?/ what am I
InnoFSPEC
/?/ ls -lrt
2012 raman spectroscopy
2012 p3d data reduction
2013 testing optical fibre bundles
2013 erasmus and virus spectrographs
2014 frequency comb
    
```



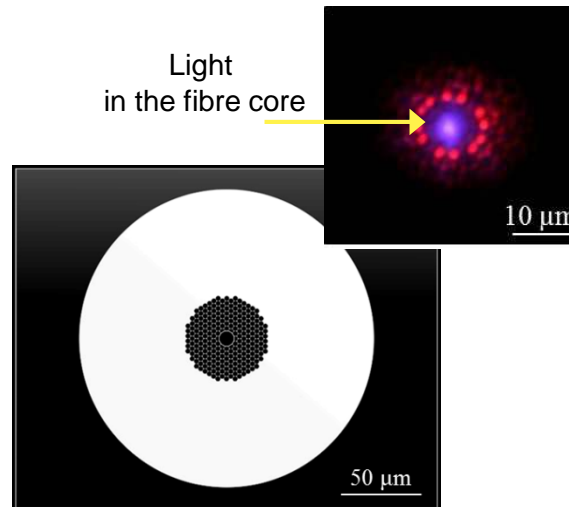
Output of the fibre
(supercontinuum)



Highly-Nonlinear Photonic Crystal Fibre

Applied for supercontinuum source
(it is a broad spectrum 350nm – 1700 nm)

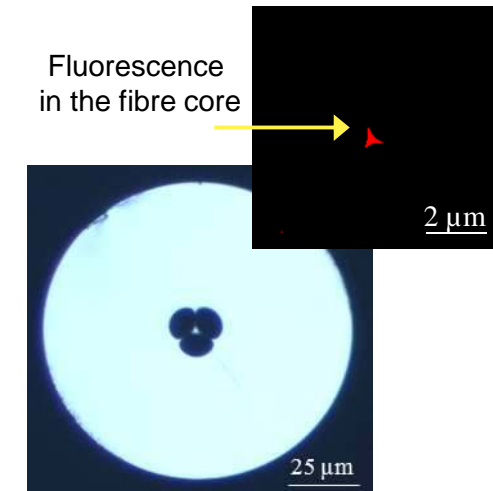
Light
in the fibre core



Hollow-Core Photonic Crystal Fibre

Applied for sensors: DNA and liquid viscosity

Fluorescence
in the fibre core



Suspended-Core Optical Fibre

The AIP story begins with.....

1. Multichannel Raman Spectroscopy (MRS)

The first results for combining:

1. Erasmus-F Spectrograph
2. Fibre Bundle (400 optical fibres)
3. Microscope objectives and filters
4. Organic and non-organic samples

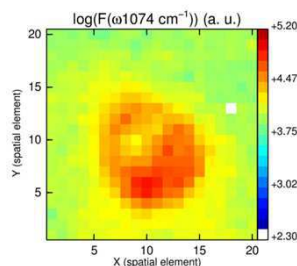


Fig. Map of a pork sample

Because of MRS.....

2. Data reduction using p3d

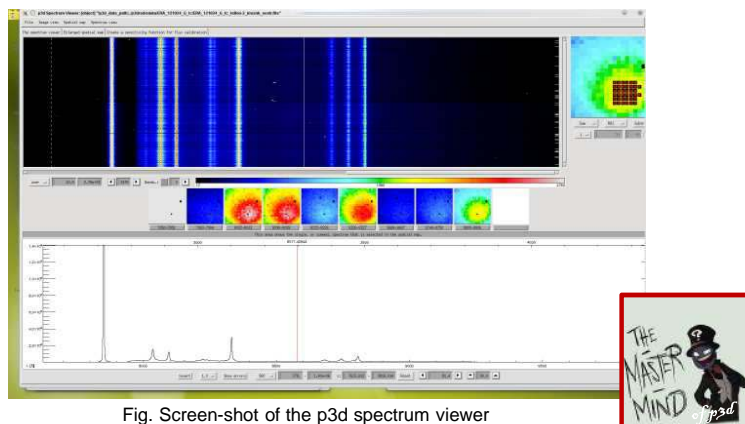


Fig. Screen-shot of the p3d spectrum viewer

Cool, now you work with....

3. Data reduction and analysis for:

Optical fibre bundles connected to:

1. Erasmus-F Spectrograph
2. VIRUS Spectrograph

And, as well....

4. Frequency comb

1. Testing the frequency comb with Erasmus-F
2. Preparing the optical setup for transport to Calar Alto and to Potsdam
3. Data reduction and analysis

But...sometimes there are surprises in the story:

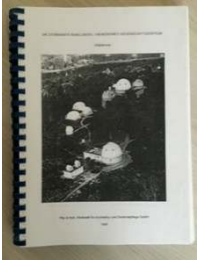
as attending the first sky-test of the frequency comb in Calar Alto Observatory, Spain (!)



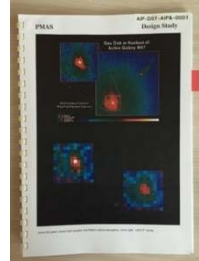
....to be continued....

Martin M. Roth
innoFSPEC

Zielplanung AIP (1995)



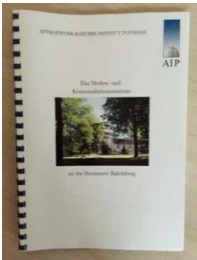
PMAS (1996)



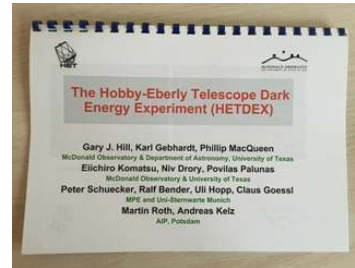
innoFSPEC (2008)



MCC (1999)



VIRUS (2002)



3Dsensation (2013)



MUSE (2003)



Multiplex Raman (2012)



MUSE GTO
NGC300

g f e h a b c d q i k j p l m n o

[illegible]

GTO-02-P94

094.D-0116(A)	NGC300-B	●	6	1.4h	1.1h			1.1h						0.83 – 0.96	1.13 – 16.36
094.D-0116(A)	NGC300-C	●	9	2.2h	1.6h	1.6h								0.57 – 0.72	-7.77 – 48.20
094.D-0116(A)	NGC300-I	●	3	41.4m	32.0m	32.0m								0.50 – 0.51	15.65 – 22.85
094.D-0116(A)	All		18	4.3h	3.2h	2.1h	1.1h								
094.D-0116(A)	A+B only			4.3h	3.2h										

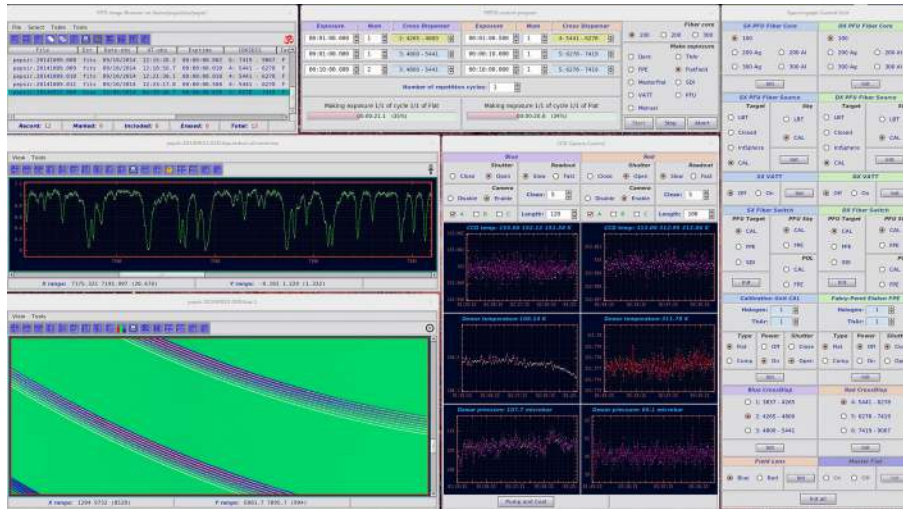
GTO-02-P94															
094.D-0116(A)	NGC300-B	●	6	1.4h	1.1h			1.1h						0.83 – 0.96	1.13 – 16.36
094.D-0116(A)	NGC300-C	●	9	2.2h	1.6h	1.6h								0.57 – 0.72	-7.77 – 48.20
094.D-0116(A)	NGC300-I	●	3	41.4m	32.0m	32.0m								0.50 – 0.51	15.65 – 22.85
094.D-0116(A)	All		18	4.3h	3.2h	2.1h	1.1h								
094.D-0116(A)	A+B only			4.3h	3.2h										

GTO-02-P94

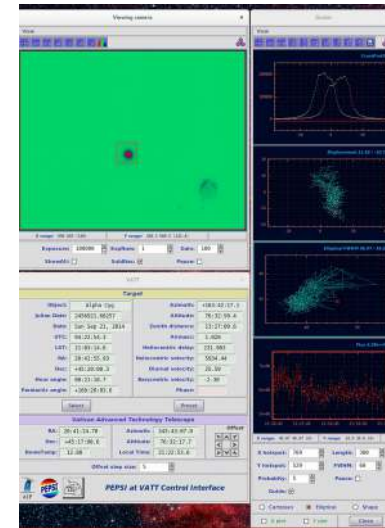
094.D-0116(A)	NGC300-B	●	6	1.4h	1.1h			1.1h						0.83 – 0.96	1.13 – 16.36
094.D-0116(A)	NGC300-C	●	9	2.2h	1.6h	1.6h								0.57 – 0.72	-7.77 – 48.20
094.D-0116(A)	NGC300-I	●	3	41.4m	32.0m	32.0m								0.50 – 0.51	15.65 – 22.85
094.D-0116(A)	All		18	4.3h	3.2h	2.1h	1.1h								
094.D-0116(A)	A+B only			4.3h	3.2h										

Ilya Ilyin
High-res Spectroscopy

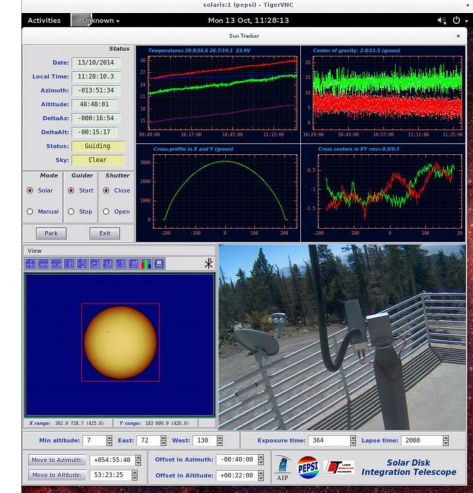
S4S is an object-oriented software is used to control various components of PEPSI hardware and for the data reduction of echelle images and analysis of high-resolution integral light and polarized spectra.



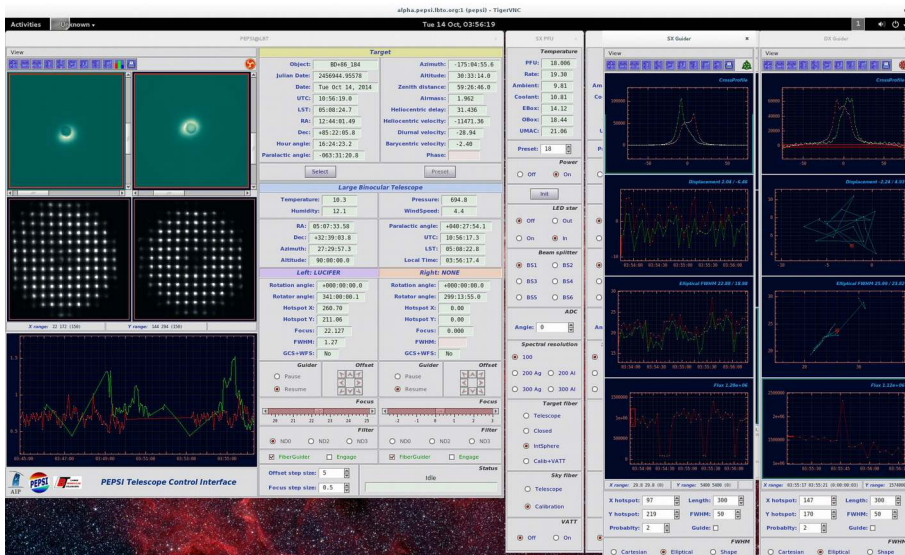
Main Control Interface of PEPSI spectrograph with all subsystems attached.



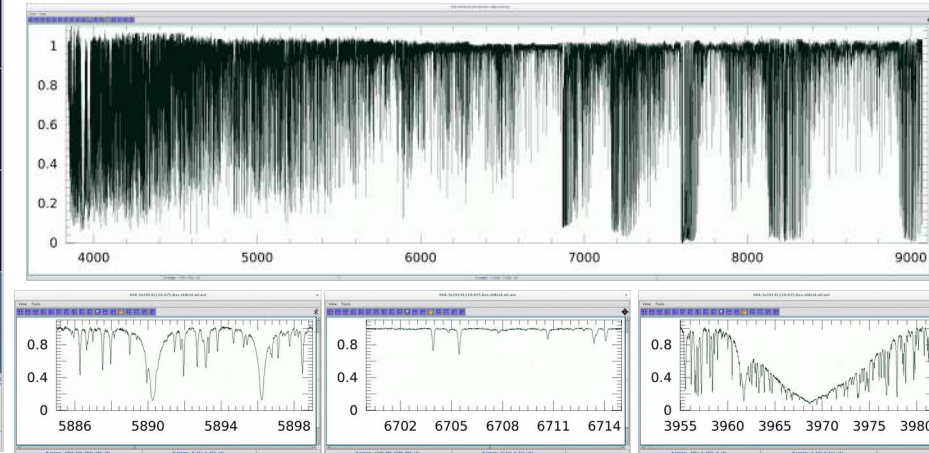
Control interface to Vatican Advanced Telescope from PEPSI with guiding on a target shown



Solar Disk Integration Telescope Control Interface allows to guide of the Sun and make exposures with PEPSI in automatic mode



Guiding with PFU at LBT with both SX and DX focal stations



The first light Solar spectrum at R=270000

Gabriele Schönherr

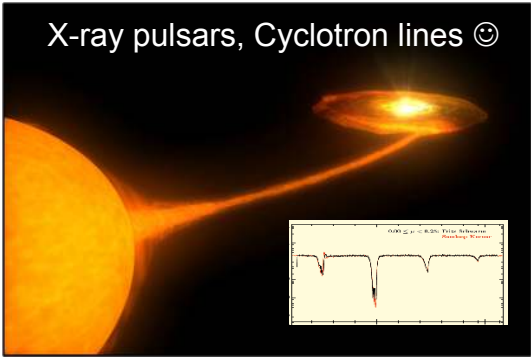
Vorstandsreferentin / Galaxies

Projects - my scientific and my non-scientific me

Scientist...



X-ray astronomy group



X-ray pulsars, Cyclotron lines ☺



Bamberg, Tübingen, Madrid

Science Communicator...



Science Manager...



...many emails...

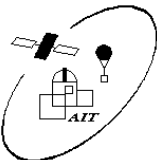
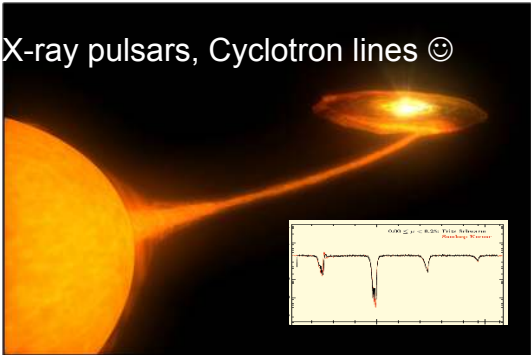


Projects - my scientific and non scientific me

Scientist...



X-ray astronomy group



Bamberg, Tübingen, Madrid

Science Communicator...

Life possible on extrasolar moons

10 January 2013. In their search for habitable worlds, astronomers have started to look for planets outside the solar system. In a new study, a pair of researchers has found that moons could be habitable too.

Where are all the dwarfs?

1 February 2013. Astronomers of the International Dark Energy Survey (IDES) are explaining the famous missing dwarf problem: the Cold Dark Matter and Dark Energy.

THE UNIVERSE YOURS TO DISCOVER

BEYOND INTERNATIONAL YEAR OF ASTRONOMY



Science Manager...



Emails...



February 2015 +



No more SAB/Evaluation E-mails!!!



New Science Communications Office, Berlin

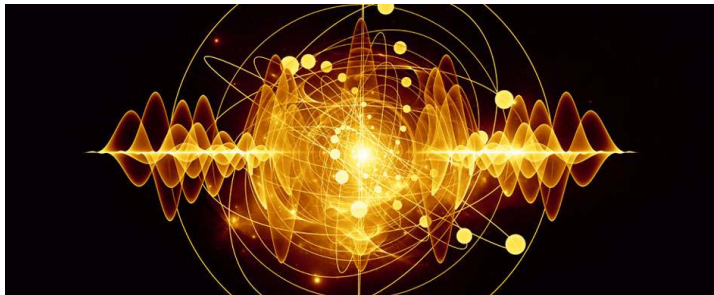
February 2015 +



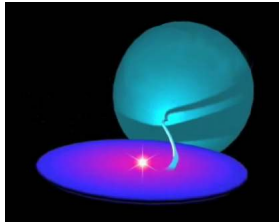
AIP

No more SAB/Evaluation E-mails!!!

Leaving AIP? Only partly. Still around for:



Communication projects: UN Year of Light 2015



Science projects, currently: Hercules X-1, X-ray + optical with STELLA?



New Science Communications Office, Berlin



Happy to discuss more after
AIP Evaluation!! 😊